



European Solar Telescope

A large solar telescope
for the XXI century

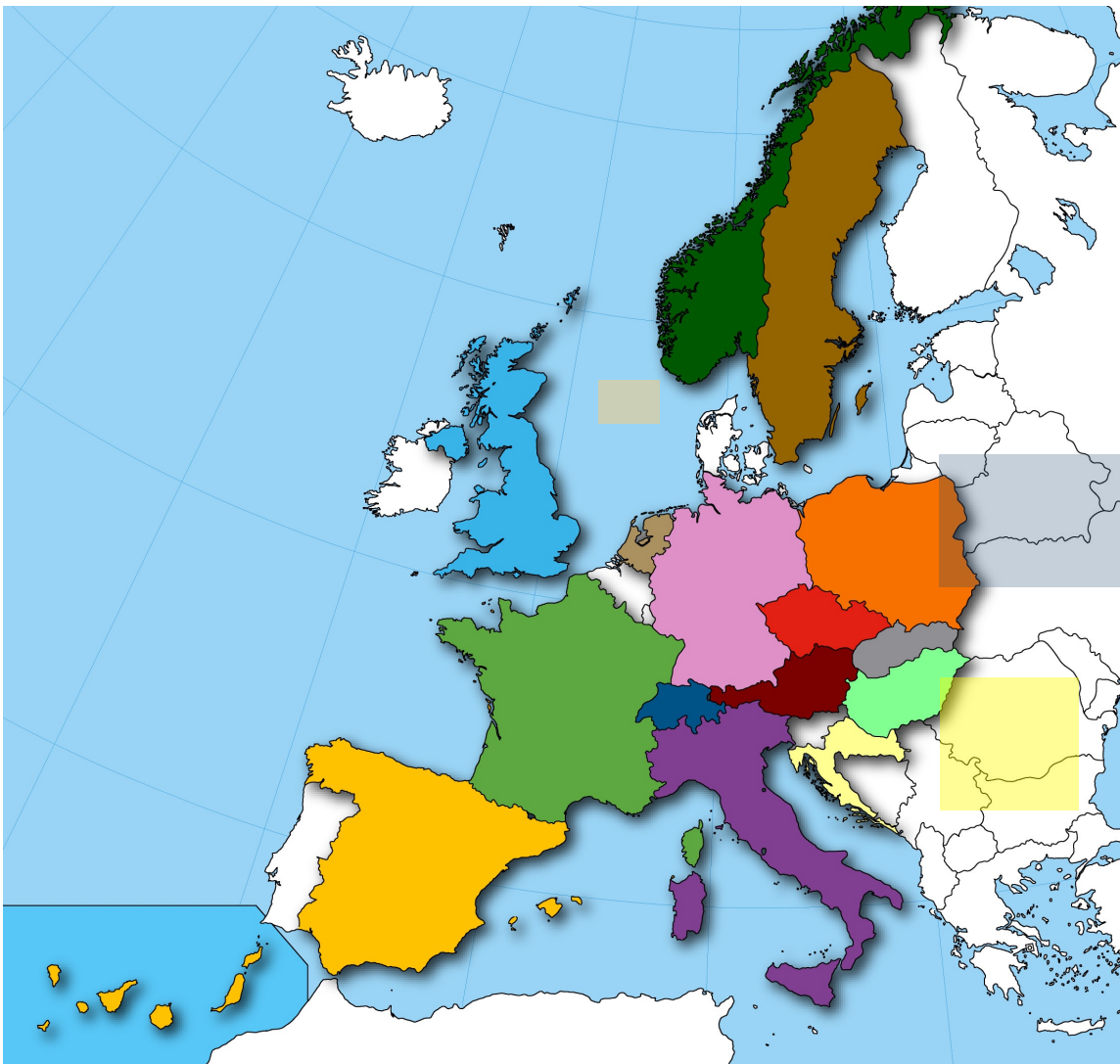


Ilaria Ermolli

INAF Osservatorio Astronomico di Roma

on behalf of the EST team

PI: Manuel Collados Vera, IAC, Spain



The EST is promoted
by the



European Association
for Solar Telescopes



EST: Strategic European Infrastructure since 03/2016



Eu-funded projects

2008-2011 Conceptual Design -- 3.2 M€

2013-2017 **SOLARNET** 

The SOLARNET logo features the word 'SOLARNET' in a bold, blue, sans-serif font. To the right of the text is a stylized graphic consisting of a blue arc and a yellow circle, resembling a rising sun or a satellite dish.

For a better and wider usage of present EU facilities to get prepared for EST -- 6 M€

2015-2018 **GREST** 

The GREST logo features the word 'GREST' in a bold, blue, sans-serif font. To the right of the text is a stylized graphic consisting of a blue arc and a yellow circle, resembling a rising sun or a satellite dish.

Continue with the technical development of EST instrumentation -- 3 M€

2017-2021 **PRE-EST** 

The PRE-EST logo features the word 'PRE-EST' in a bold, blue, sans-serif font. To the right of the text is a stylized graphic consisting of a blue arc and a yellow circle, resembling a rising sun or a satellite dish.

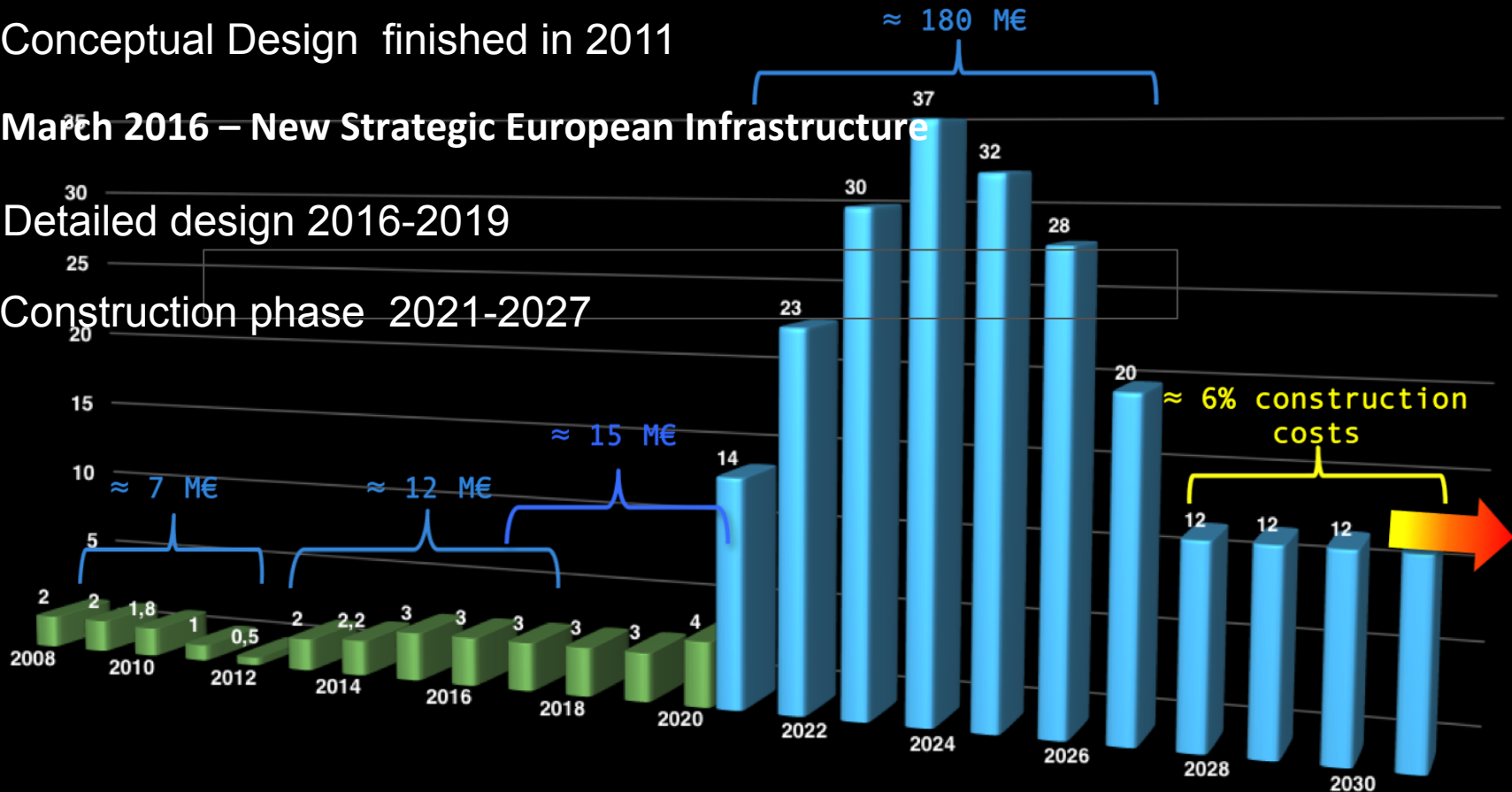
To provide the EST international consortium and the national agencies with a detailed plan for the implantation of EST -- 9 M€

Conceptual Design finished in 2011

March 2016 – New Strategic European Infrastructure

Detailed design 2016-2019

Construction phase 2021-2027



PHASE





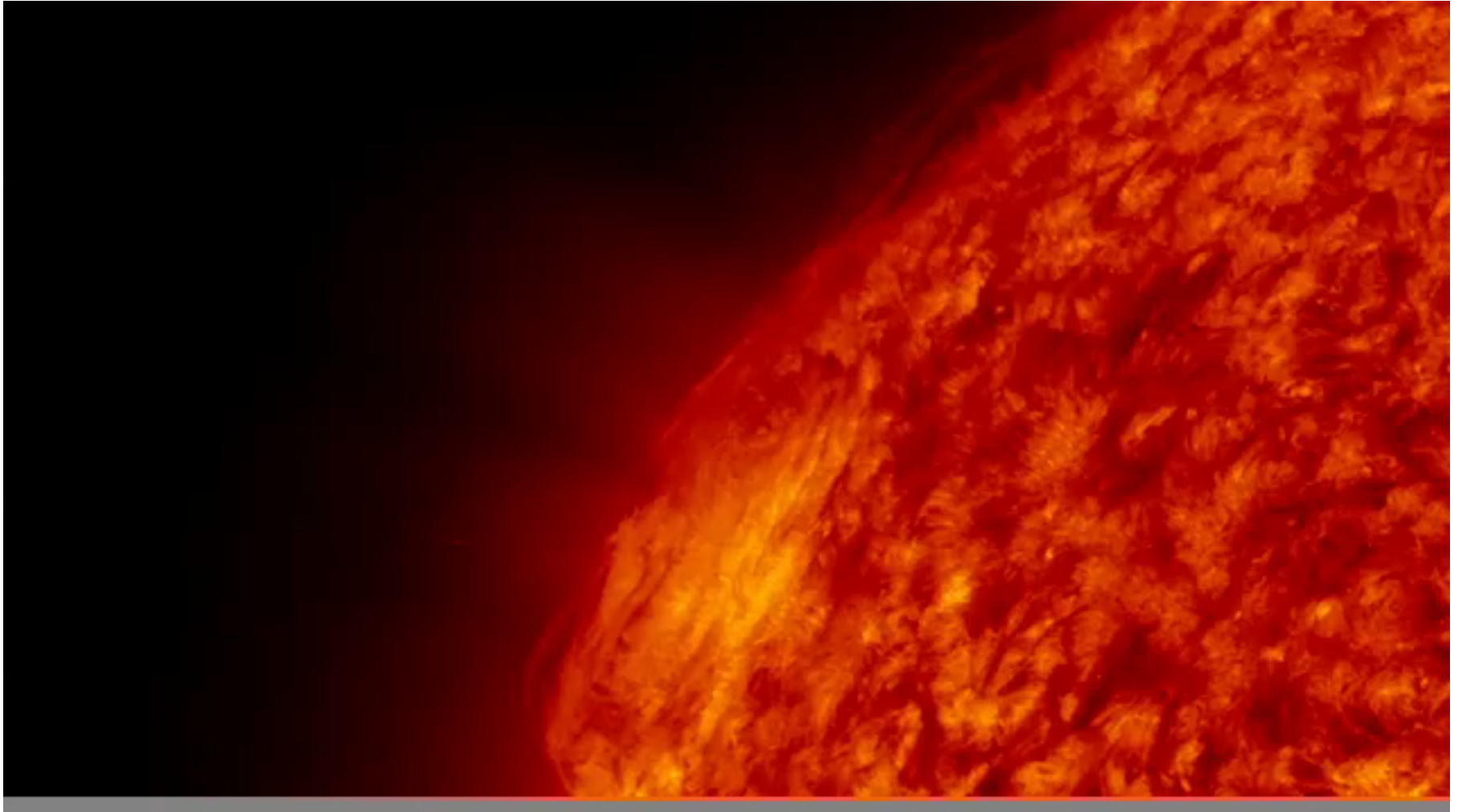
Main science goals

Increasing the capabilities of ground-based solar telescopes will allow us to provide an answer to the following questions:

- How does the **magnetic field evolve** and emerge to the solar surface?
- How is the **energy transported** from the photosphere to the chromosphere?
- How is the **energy released** and deposited in the upper atmosphere?
- Why does the Sun have a **hot chromosphere** and a hot **corona**?
- What causes the **explosive events**, flares, filament eruptions, CMEs?



Key questions



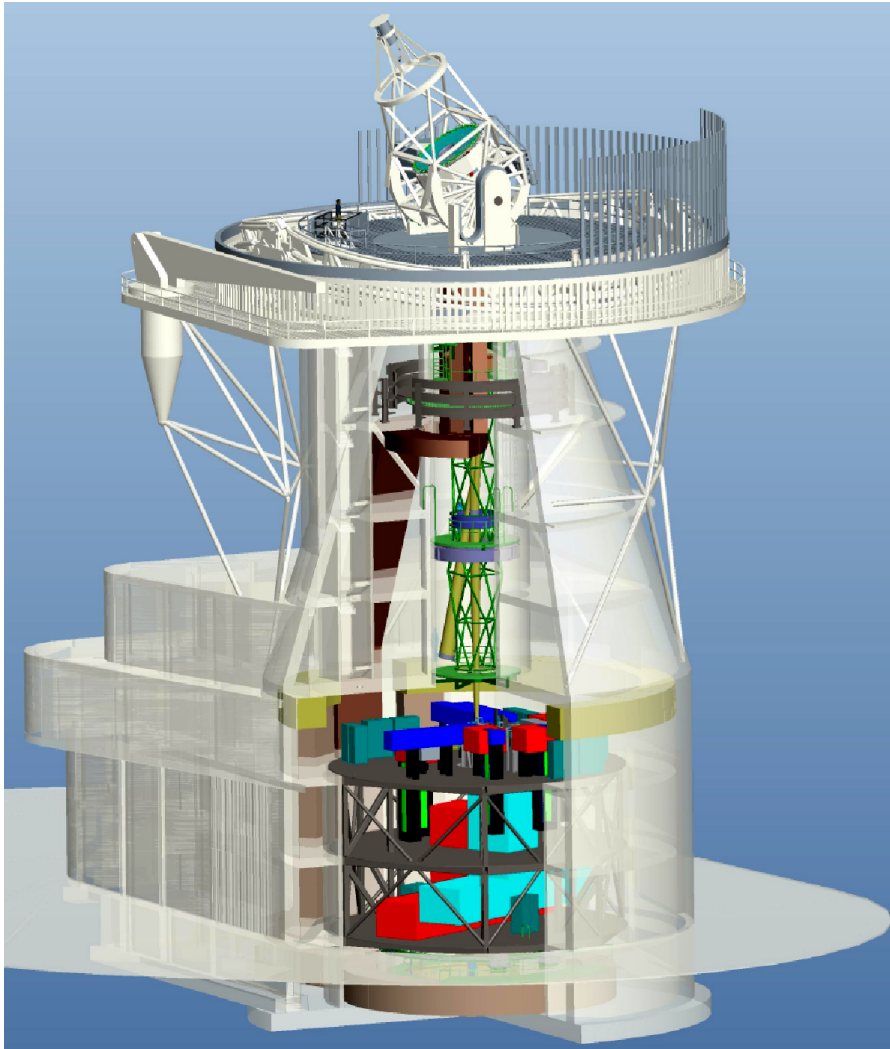


Telescope and instrumentation key requirements

- **High precision polarimetric capabilities** for accurate magnetic field determination
- **Multiwavelength capabilities** for simultaneous observation of different heights in the solar atmosphere
- **Large photon collector**
- **High angular resolution**



Conceptual design



4-m telescope to collect photons

On axis Gregory

Mirrors polarimetrically compensated
for high accuracy spectropolarimetry

**Adaptive Optics and Multi Conjugate
Adaptive Optics system** for the
highest spatial resolution

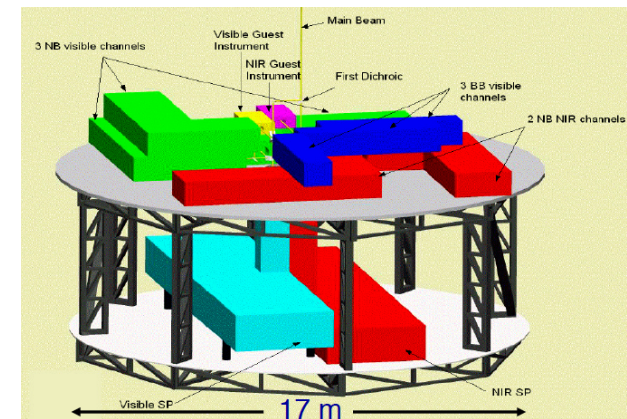
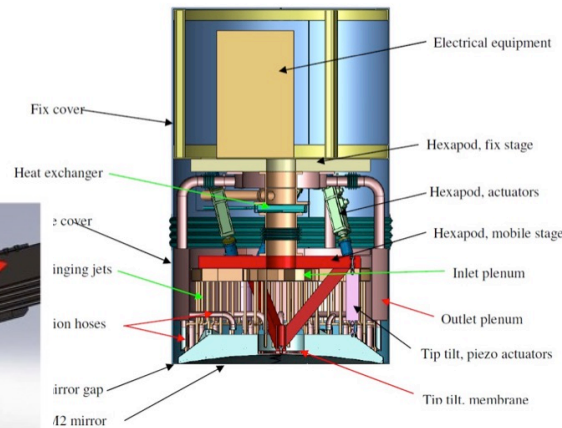
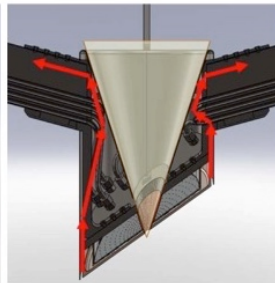
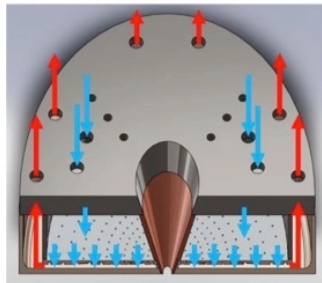
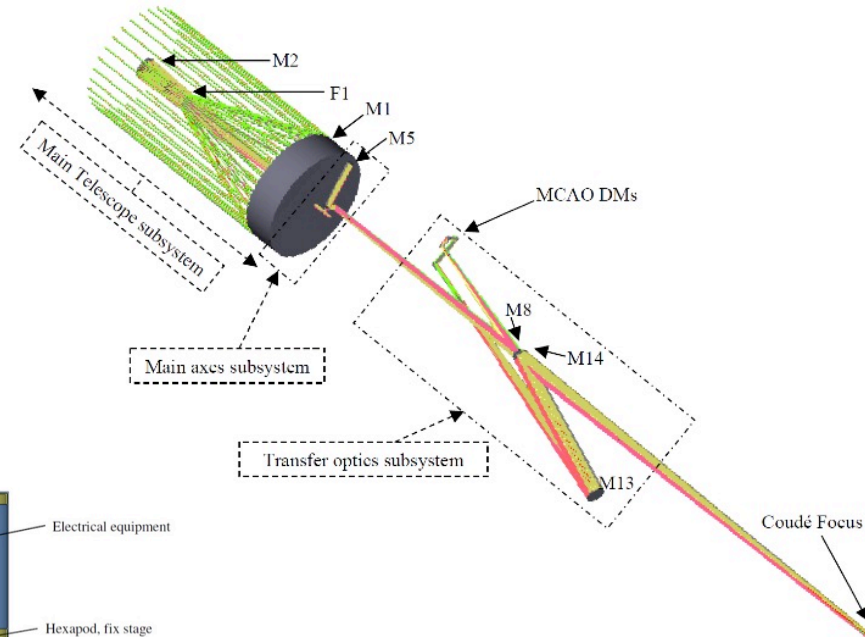
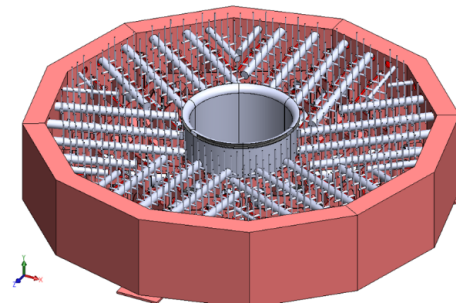
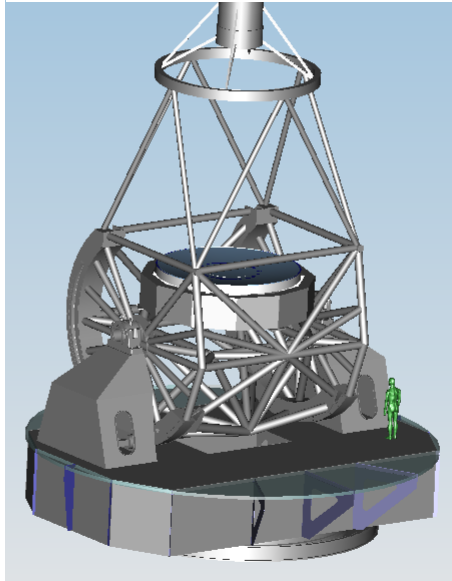
MCAO integrated in the optical path

Simultaneous instrument stations
(each with several wavelength
channels)

Coudé instruments for larger stability



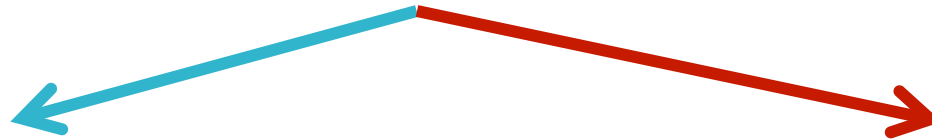
Telescope conceptual design





Instruments conceptual design

The light beam is **divided** into **infrared** and **visible** light



broad band imager

380-500 nm 380-500 nm 600-900 nm

narrow band imager

390-500 nm 500-620 nm 620-860 nm

grating spectrographs

390-560 nm 560-1100 nm

narrow band imager

800-1100 nm 500-1800 nm

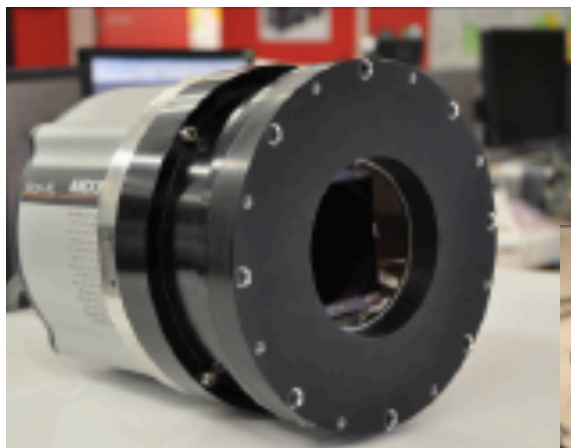
grating spectrographs

700-1600 nm 1000-2300 nm

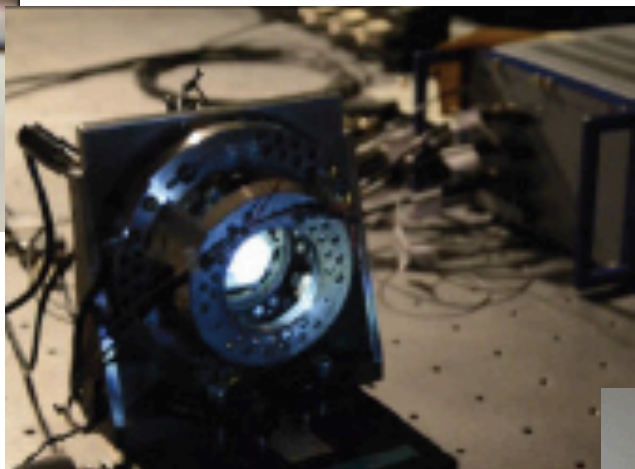
guest instrument



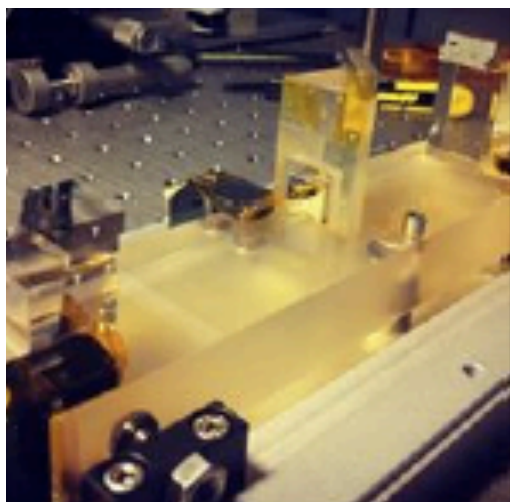
Instrumental developments



Boosting the new generation of detectors
Large format, high precision, low noise



Development of a large Fabry-Pérot prototype for high mechanical stability and high parallelism of the etalons



New techniques for 2-D spectropolarimetry
Multi-slit
Integral field unit



Development of large format liquid crystal modulators

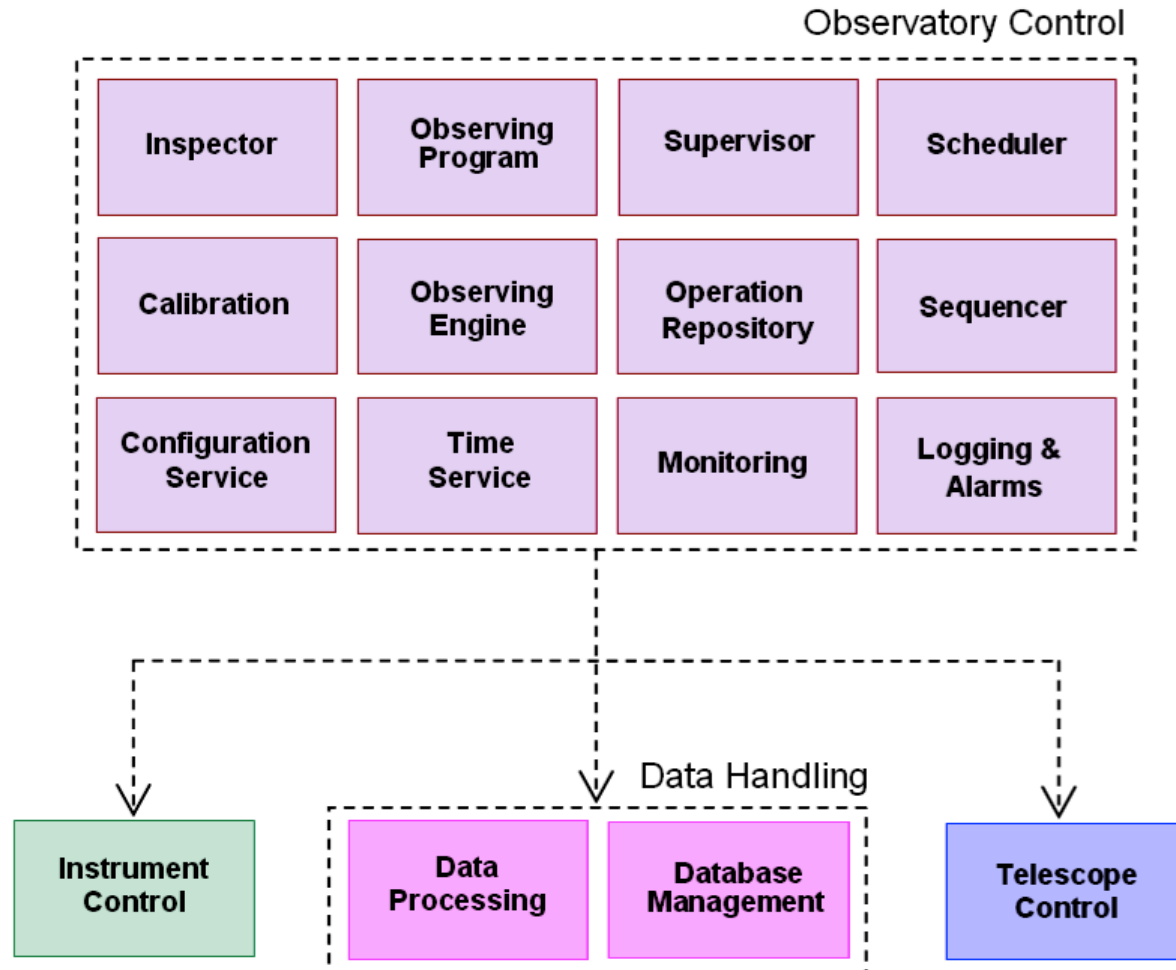


Control conceptual design

- **Control system:**
Integrated control system for the entire observatory

- Common software
- Distributed, object-oriented architecture

- **Data handling**





DH conceptual design

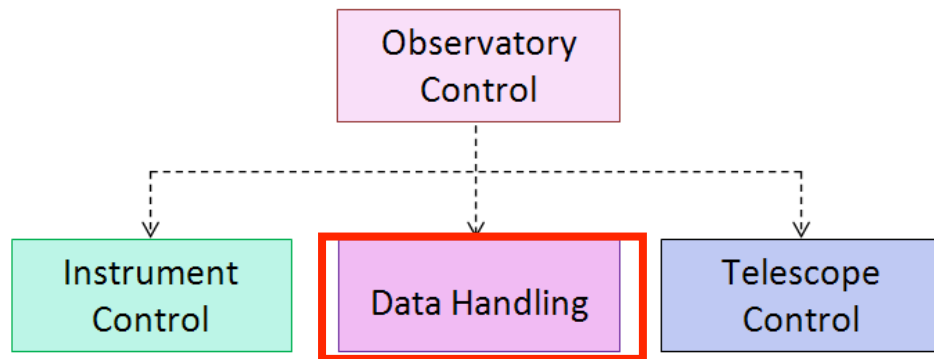
Maximum Detector Flux:

4096 x 4096 pixels ; 100 frames /sec

-> **3 GB /sec (visible)**

2048 x 2048 pixels ; 100 frames /sec

-> **0.8 GB /sec (infrared)**



Instrument	# of Detectors	Data Rate
Broadband Filters	9 visible	28 GB / sec
Narrowband Filters	9 visible 6 infrared	28 GB / sec 5 GB / sec
Spectrographs	5 visible 3 infrared	16 GB / sec 3 GB / sec
	32 detectors	80 GB / sec

On-site

Off-site

A lot & urgent

**Science-ready
VO-compliant**

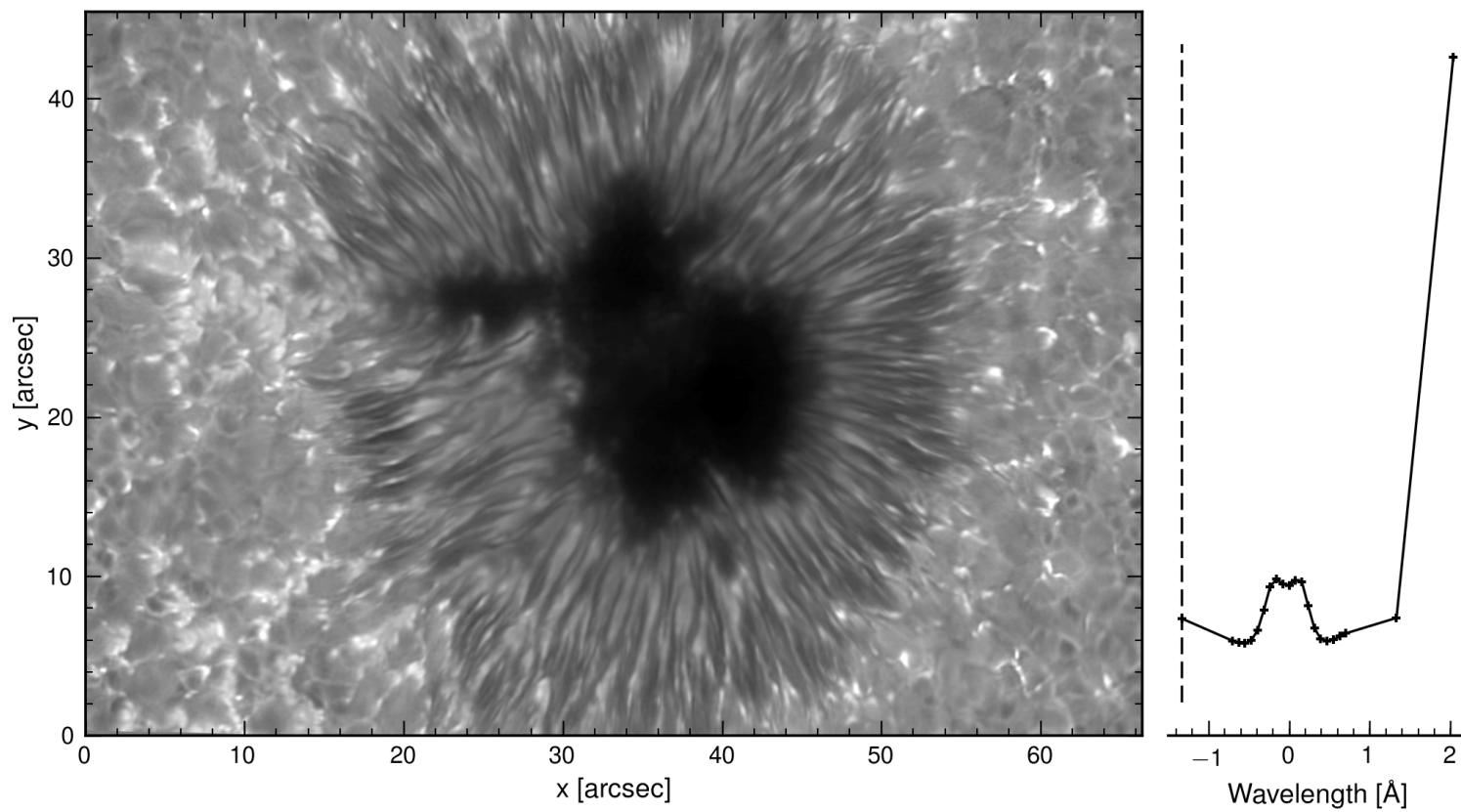
Lossless compression
Image reconstruction

Definition of standards
Pipelines
Analysis Tools



EST data

Narrow band spectropolarimetric imager

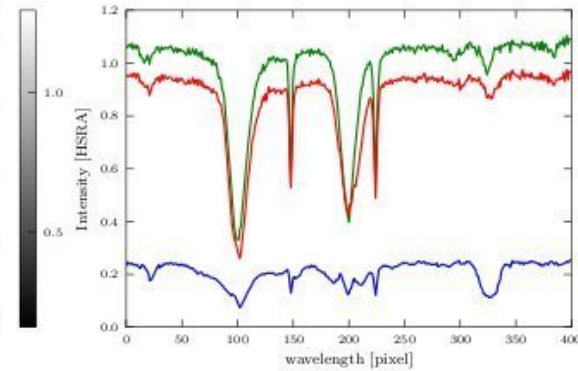
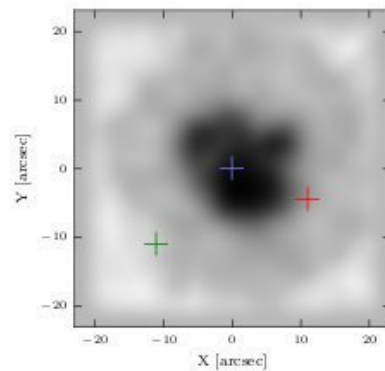
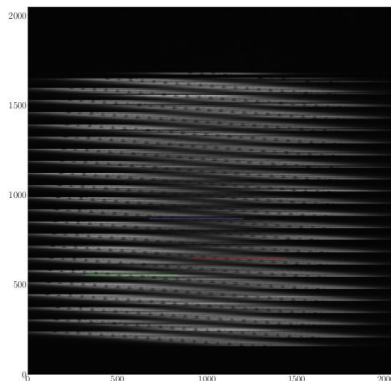
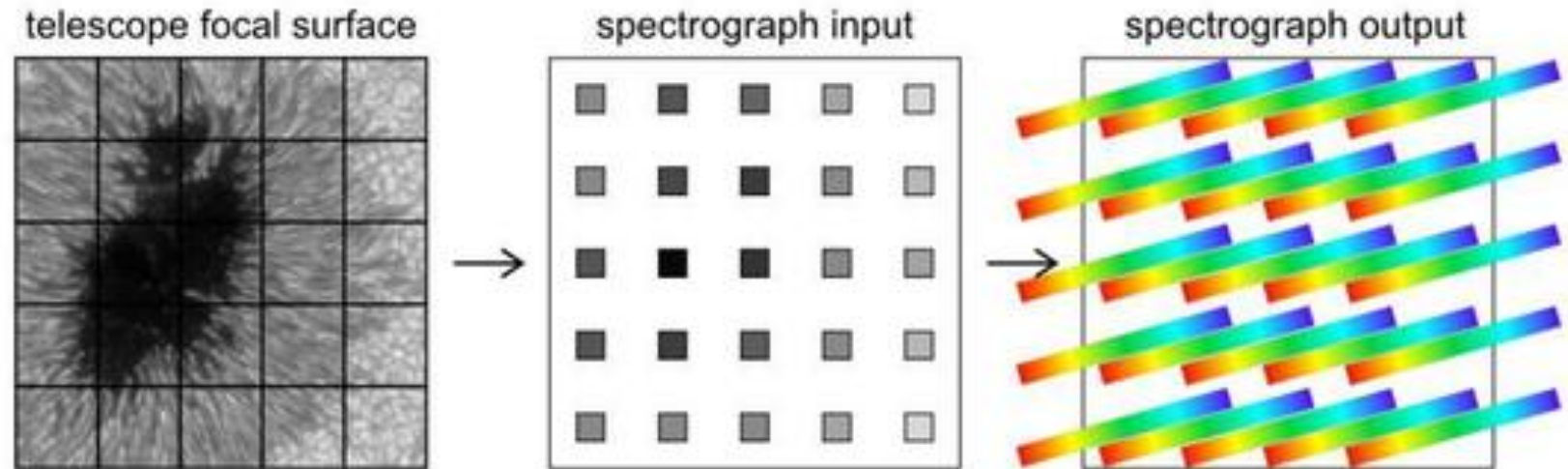


SST/CHROMIS observations: courtesy G. Scharmer



EST data

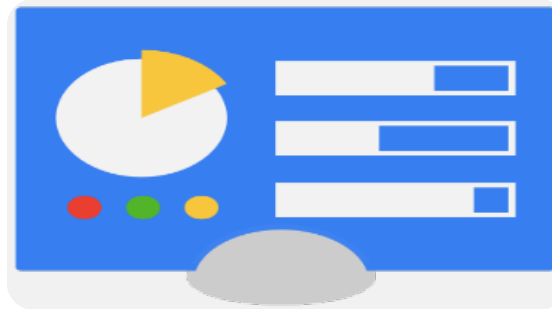
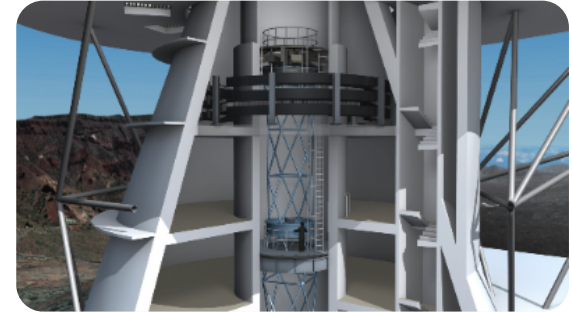
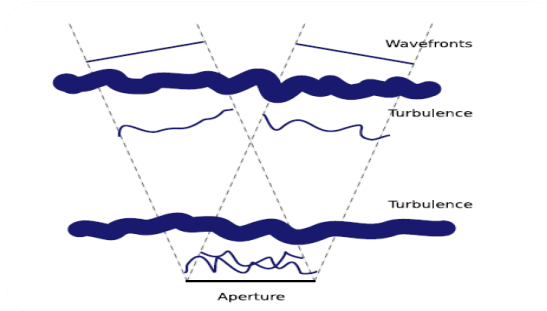
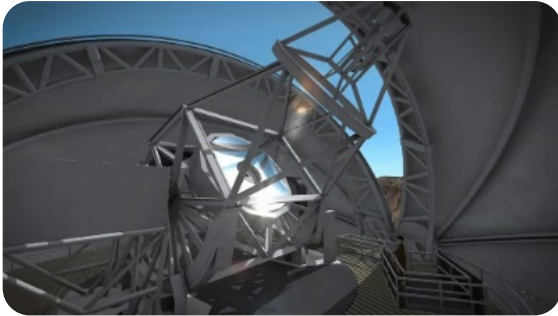
Microlens-fed spectrograph



SST/MIHi observations: courtesy M. Van Noort



Preparatory phase Technical works



2017-2021

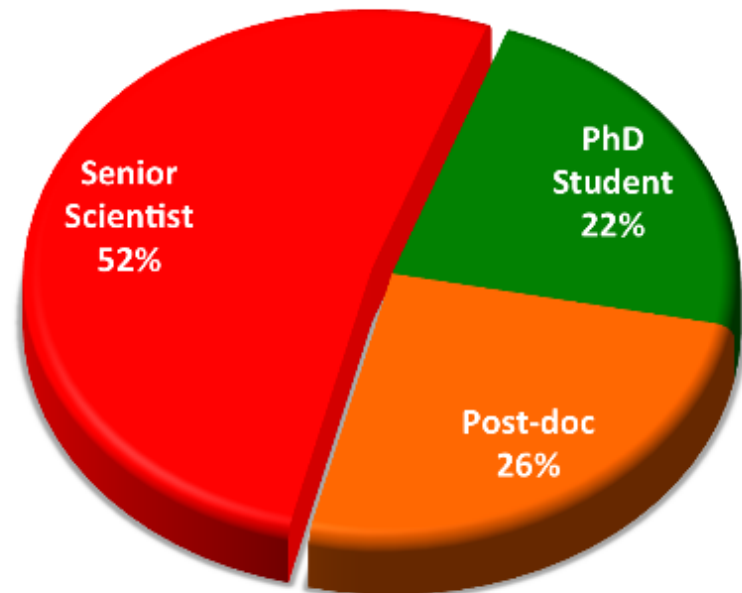
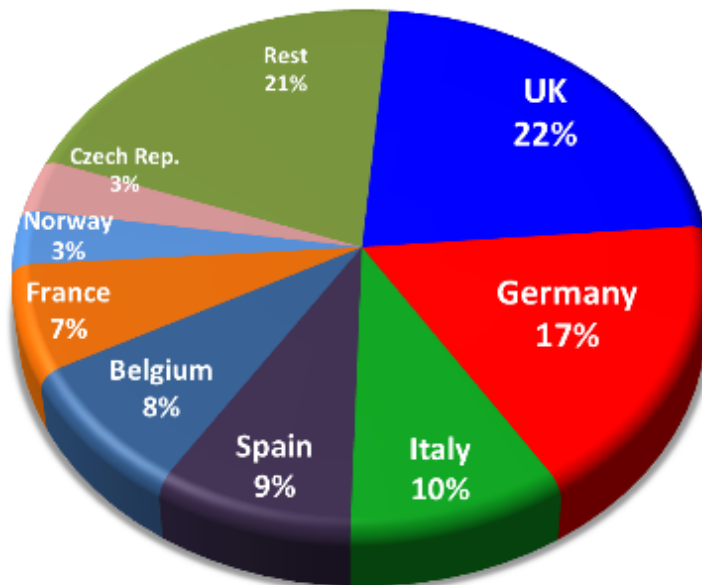


Governance, Legal Entity, Financial plan,
Site.



EST Community

1. A total of **615 researchers** have been identified in **22 European countries**.
2. **UK, Germany and Italy** represent **close to 50%** of the total solar community.
3. The first **9 countries** concentrate **more than 80%** of the community.





Thanks!

