

# CTA DADI Status

## Cherenkov Telescope Array

Mathieu Servillat, Catherine Boisson, Julien Lefaucheur



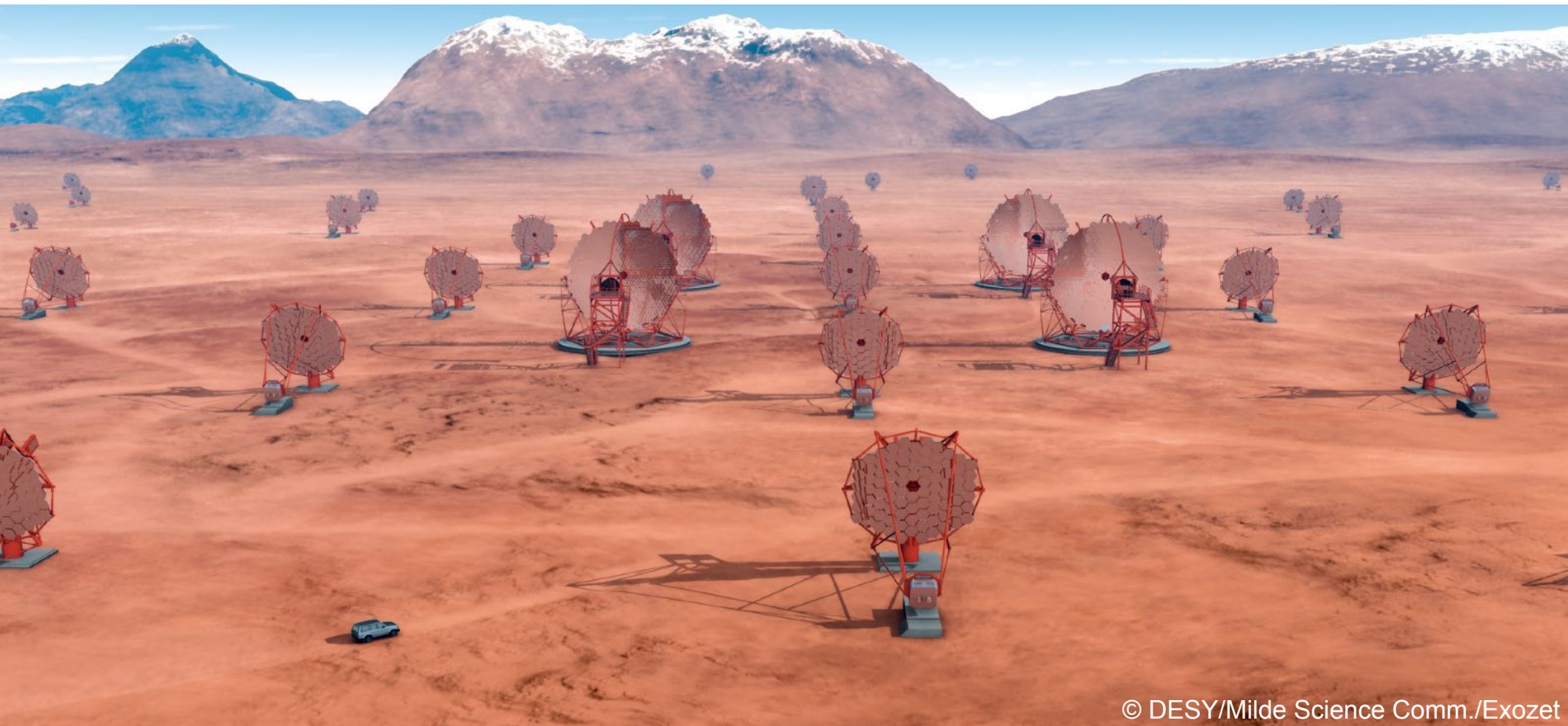
Laboratoire Univers et Théories  
Observatoire de Paris  
PSL Research University

ASTERICS DADI ESFRI Forum 2  
INAF - Osservatorio Astronomico di Trieste  
13-14 Dec. 2017





- ◆ Two arrays of 100 (South) et 20 (North) telescopes
- ◆ July 2015: sites selection, Chile (ESO) and La Palma
- ◆ 2016: pre-production phase
- ◆ 2018-2013: production phase
- ◆ Observatory open to the community

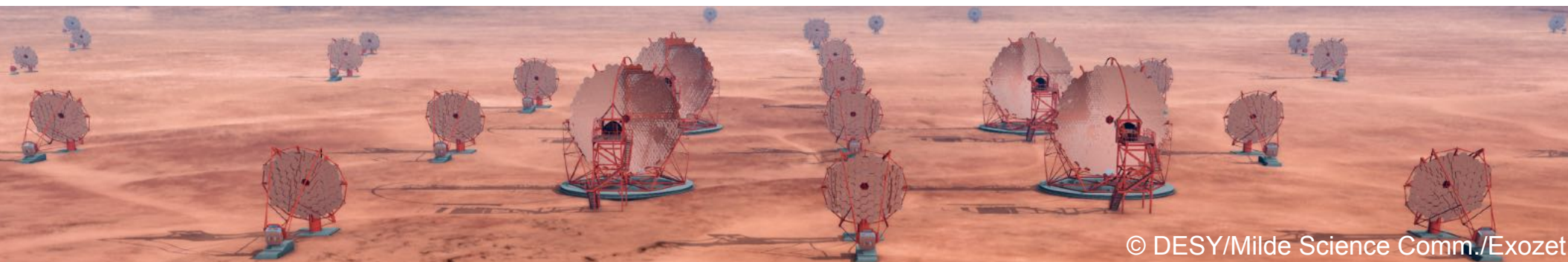
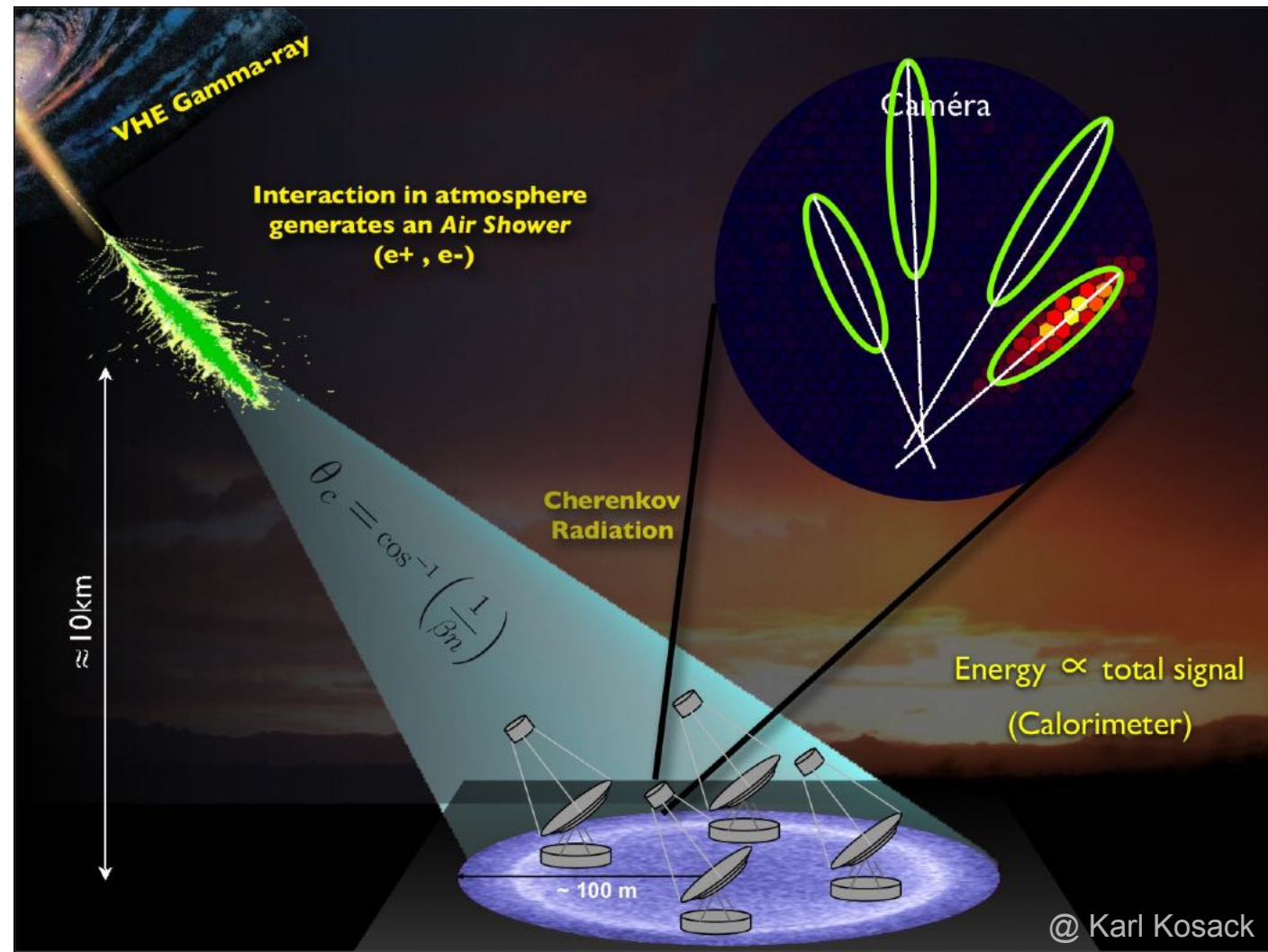


© DESY/Milde Science Comm./Exozet



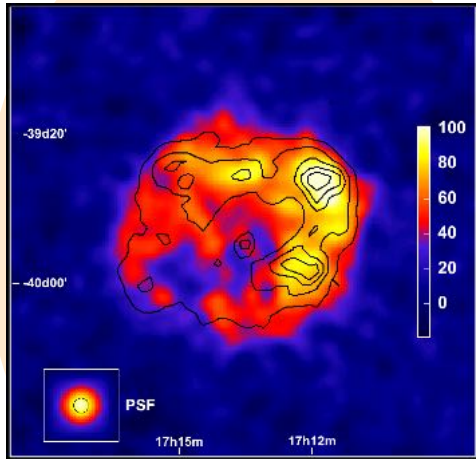
# Cherenkov Astronomy Principles

- ◆ **Dark nights** (small duty cycle)
- ◆ **Event Reconstruction:** photon, particle shower, Cherenkov light (faint, few nanoseconds)
- ◆ **Atmosphere = calorimetre**  
Simulations, assumptions
- ◆ **Complex Metadata,**  
need to be structured



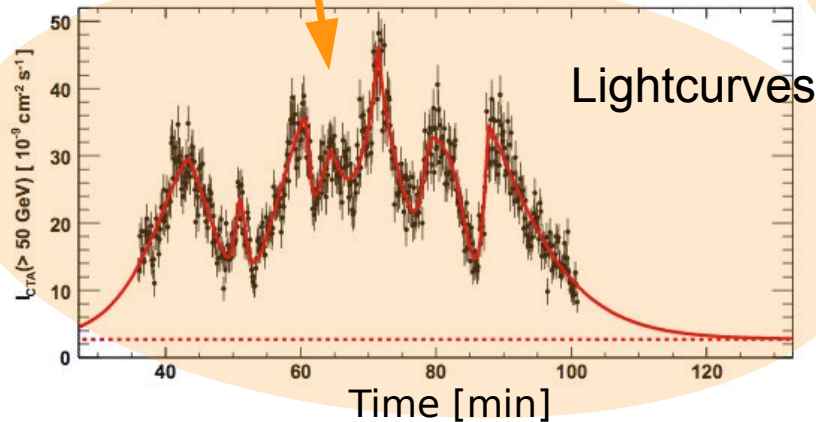
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# Multi-wavelength analysis

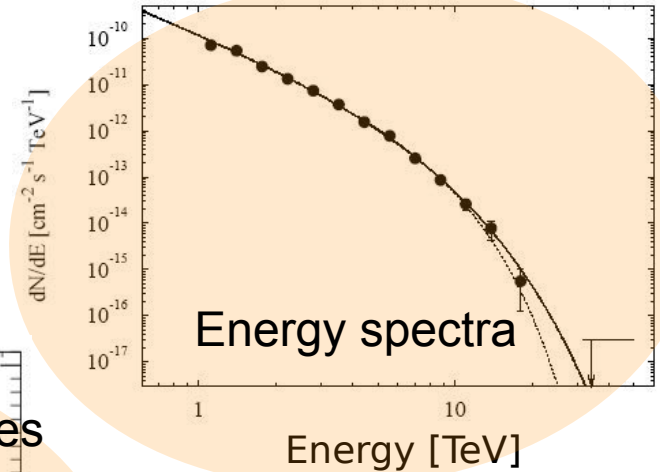


Images

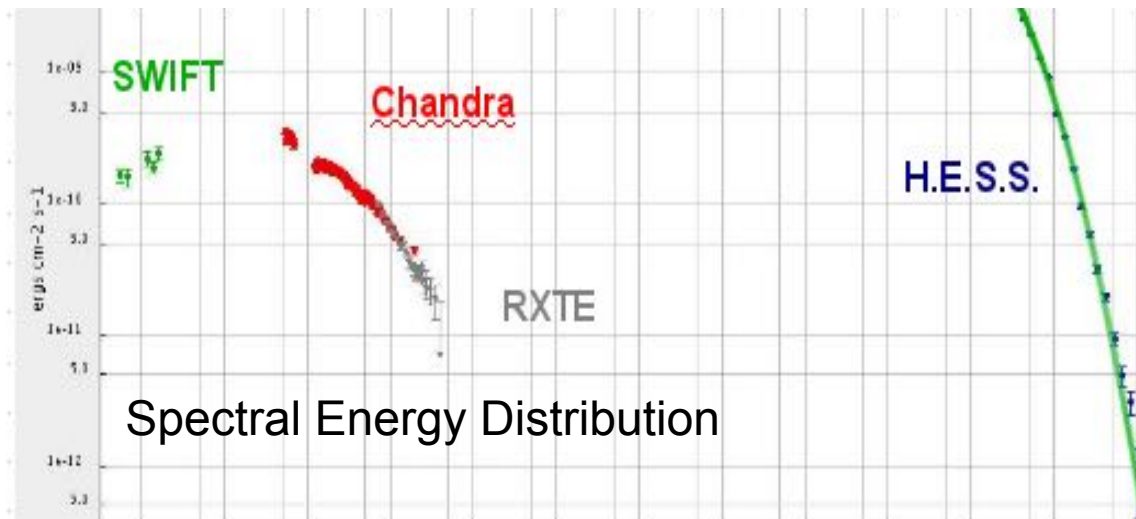
CTA Event lists  
(coordinates, time, energy)



Lightcurves



Energy spectra



Spectral Energy Distribution

Compatible data  
at other wavelength?

Simultaneous  
Calibrated  
Specific Processing?  
Context?

# VO data access prototype

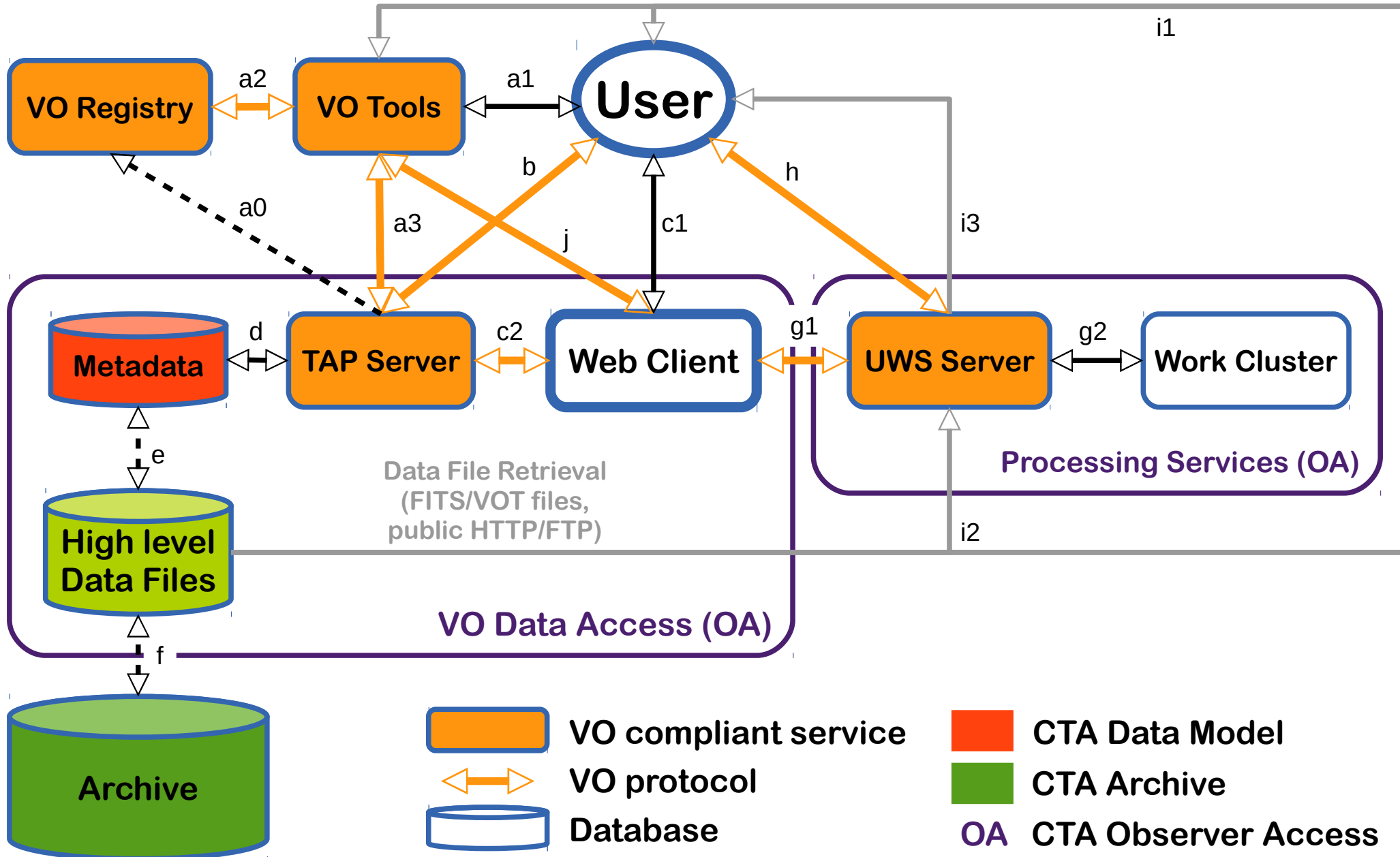
- ◆ **CTA Data Model** (not complete, still evolving)
  - ◆ [https://forge.in2p3.fr/projects/model/wiki/UML\\_models](https://forge.in2p3.fr/projects/model/wiki/UML_models)
  - ◆ Automatic Conversion **UML** to **SQL**
  - ◆ Relational database implemented (PostgreSQL)
- ◆ **Data Ingestion**: CTA First Data Challenge (1DC)
- ◆ **VO Compliant**
  - ◆ **IVOA ObsCore** Data Model
  - ◆ GAVO DaCHS server: **TAP**, **ADQL**
- ◆ **Web Client** (Django, jQuery, BootStrap)
- ◆ **Online Analysis**: **UWS**, **SAMP**
- ◆ **Single Sign On** solutions

} **ObsTAP**



► Complete solution based on VO standards/protocols

# VO data diffusion prototype





# CTA Data Distiller

<https://voparis-cta-test.obspm.fr>



CTA Data Distiller

🔍 Search Form

⚙️ Job List

✕ Sign out user

☒ Cone Search

Target Name

Crab Nebula

Used to query Simbad with Sesame and set RA/Dec.

Source RA (deg)

83.633

Source Dec (deg)

22.514

Search radius (deg)

0.001

Submit

Reset

- ◆ Django, jQuery, BootStrap3
- ◆ Name resolver  
(Simbad through Sesame)
- ◆ Builds and Sends the ADQL query

▼ ObsCore Search

proposal\_id

Proposal ID

dataprodut\_type

Nothing selected

Data product (file content) primary type

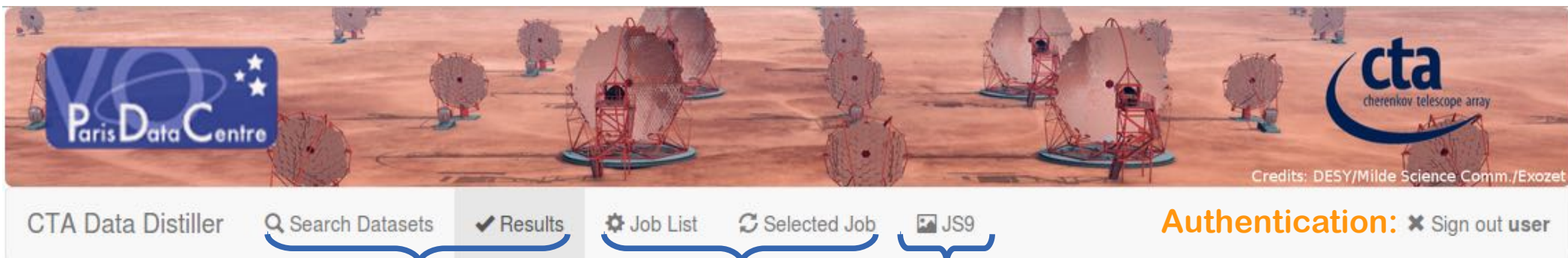
dataprodut\_level

Nothing selected

DL0-5

# CTA Data Distiller

<https://voparis-cta-test.obspm.fr>



The header of the CTA Data Distiller interface features a banner with the Paris Data Centre logo on the left and the CTA Cherenkov Telescope Array logo on the right. Below the banner is a navigation bar with several tabs: 'CTA Data Distiller', 'Search Datasets', 'Results', 'Job List', 'Selected Job', and 'JS9'. The 'Results' tab is currently selected. To the right of the navigation bar, there is an 'Authentication' section with a 'Sign out user' link.

Search

Analyse

Visualisation

Authentication: ✕ Sign out user

Results show/hide query

SELECT \* FROM cta.vo\_obscore as o WHERE 1 = intersects(o.s\_region, circle('ICRS', 83.63308333, 22.0145, 0.001))

Send

ADQL query

ObsCore fields

SAMP

Interop (SAMP)

Send Result Table

Send Selected Data

Analysis tools

Create Count Map(s)

Extract Spectrum

Plotting tools

TOPCAT

Aladin

VOSpec

SPLAT

UWS

	dataproduct_type	obs_collection	obs_id	target_name	s_ra (deg)	s_dec (deg)
<input type="checkbox"/>	eventlist	1	23592	Crab Nebula	82.01333618164062	22.01444435119629
<input type="checkbox"/>	eventlist	1	23559	Crab Nebula	85.25333404541016	22.01444435119629
<input type="checkbox"/>	eventlist	1	23526	Crab Nebula	83.63333129882812	22.51444435119629
<input type="checkbox"/>	eventlist	1	23523	Crab Nebula	83.63333129882812	21.51444435119629
<input type="checkbox"/>	eventlist	3	5003499	CrabNebula	83.28087615966797	21.784133911132812

Showing 1 to 5 of 10 rows 5 records per page

<< < 1 2 > >>



# Authentication & Authorization

Sign in through eduGAIN

OR

Sign in using CTA Unity IDM

OR

OpenID Connect



OAuth2



OAuth



OpenID 2.0

mservillat.pip.verisignlabs.com

Submit

OR

Username

admin

Password

...

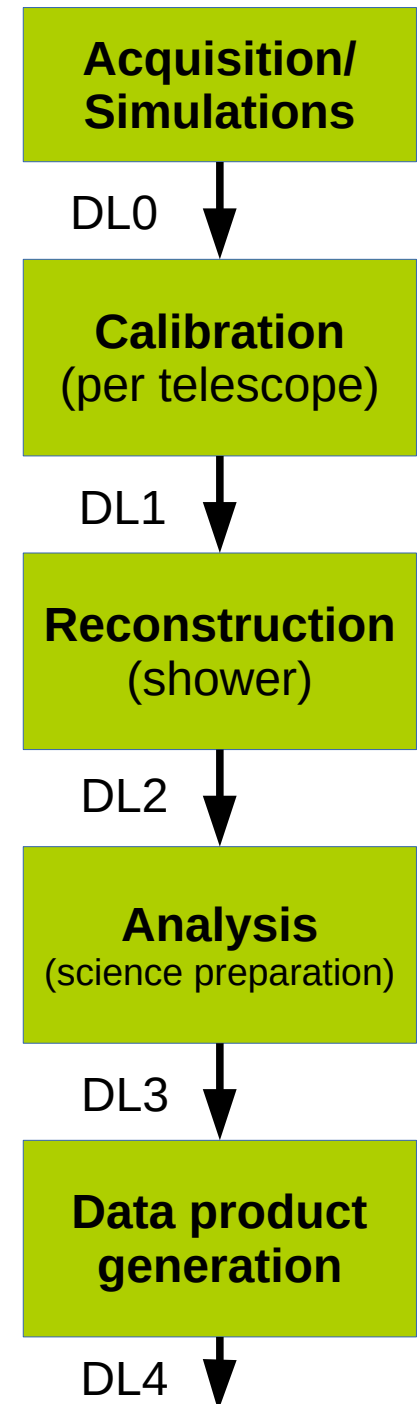
Submit

Reset

- ◆ Shibboleth+Grouper
  - ◆ EduGAIN federation
  - ◆ SAML2
- ◆ Unity IDM
  - ◆ Uses OpenID Connect
- ◆ OpenID Connect
  - ◆ Google as an IdP
- ◆ OAuth2
  - ◆ Github, Google, Facebook, ...
- ◆ OAuth
  - ◆ Twitter, ...
- ◆ OpenID 2.0 (deprecated)
- ◆ Local account

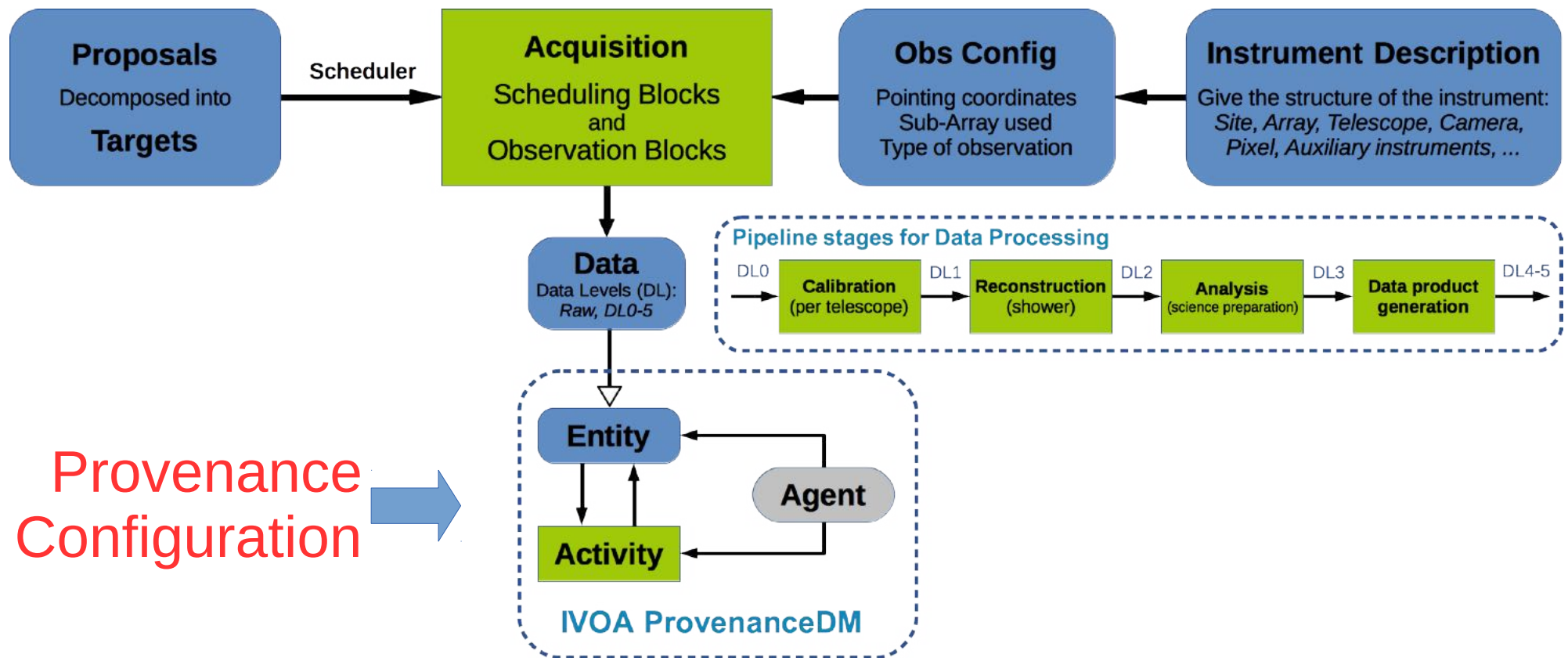
# Pipeline Requirements

- ◆ **Open** observatory
- ◆ Must ensure that data processing is **traceable** and **reproducible** (A-USER-0110)
- ◆ **Inform** user on processing steps performed
- ◆ Link to progenitor to regenerate data (DL3 to DL4)
  
- ◆ Identify how a data product was produced  
⇒ **Provenance**
- ◆ Identify what detailed options were used  
⇒ **Configuration**



# High level metadata model

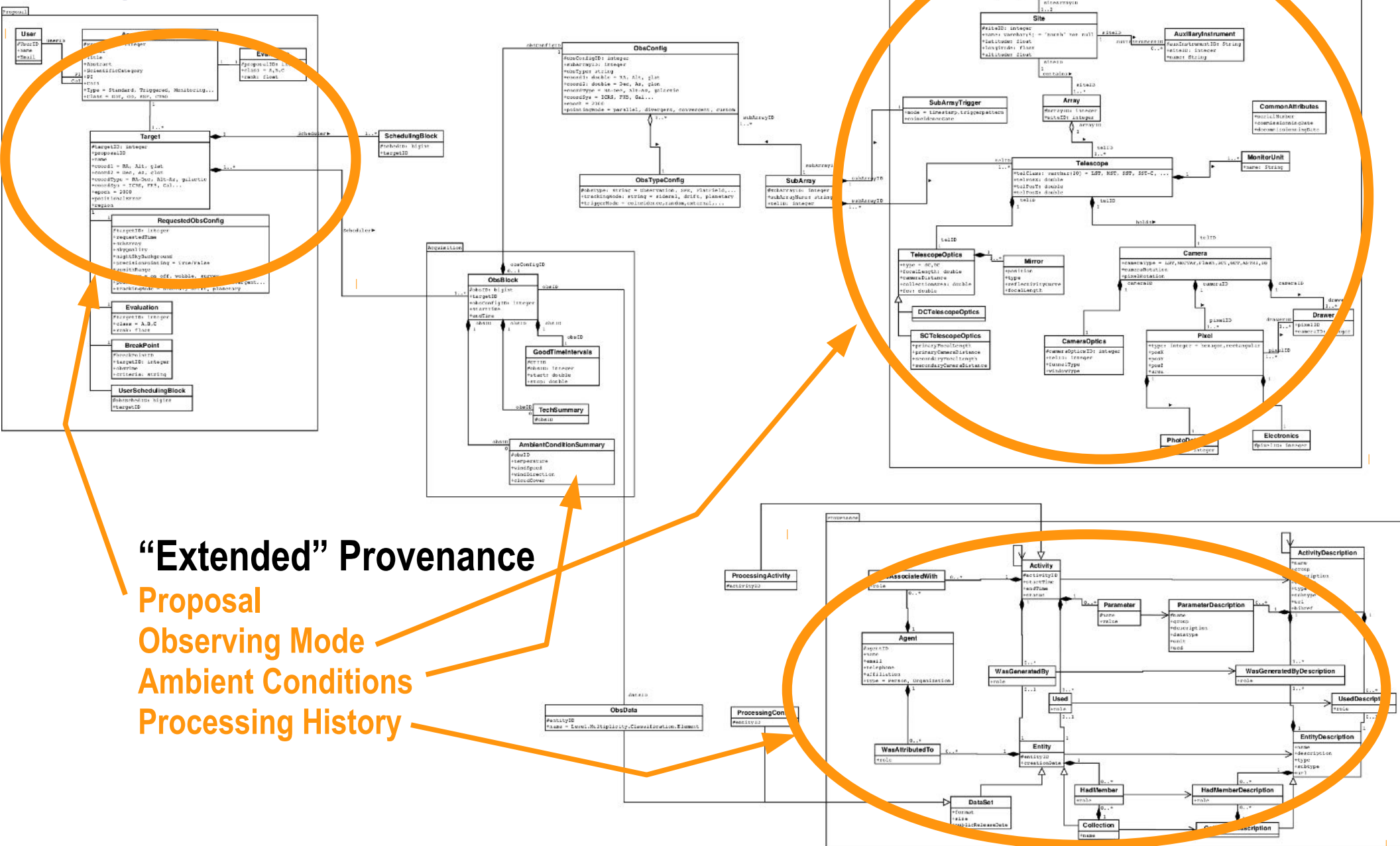
- ◆ Defines **structure** of services, content and context of data
- ◆ Can be seen as a **global interface**



Servillat et al. 2017, ADASS Trieste



# All you need is metadata!



# IVOA Provenance Data Model

## Version 1.0



### IVOA Working Draft 2017-10-12

Working group  
DM

This version

<http://www.ivoa.net/documents/ProvenanceDM/20171012>

Latest version

**<http://www.ivoa.net/documents/ProvenanceDM/>**

Previous versions

[WD-ProvenanceDM-1.0-20170921.pdf](#)

[WD-ProvenanceDM-1.0-20161121.pdf](#)

[ProvDM-0.2-20160428.pdf](#)

[ProvDM-0.1-20141008.pdf](#)

Author(s)

Kristin Riebe, Mathieu Servillat, François Bonnarel, Mireille Louys, Markus Nullmeier, Florian Rothmaier, Michèle Sanguillon,  
IVOA Data Model Working Group

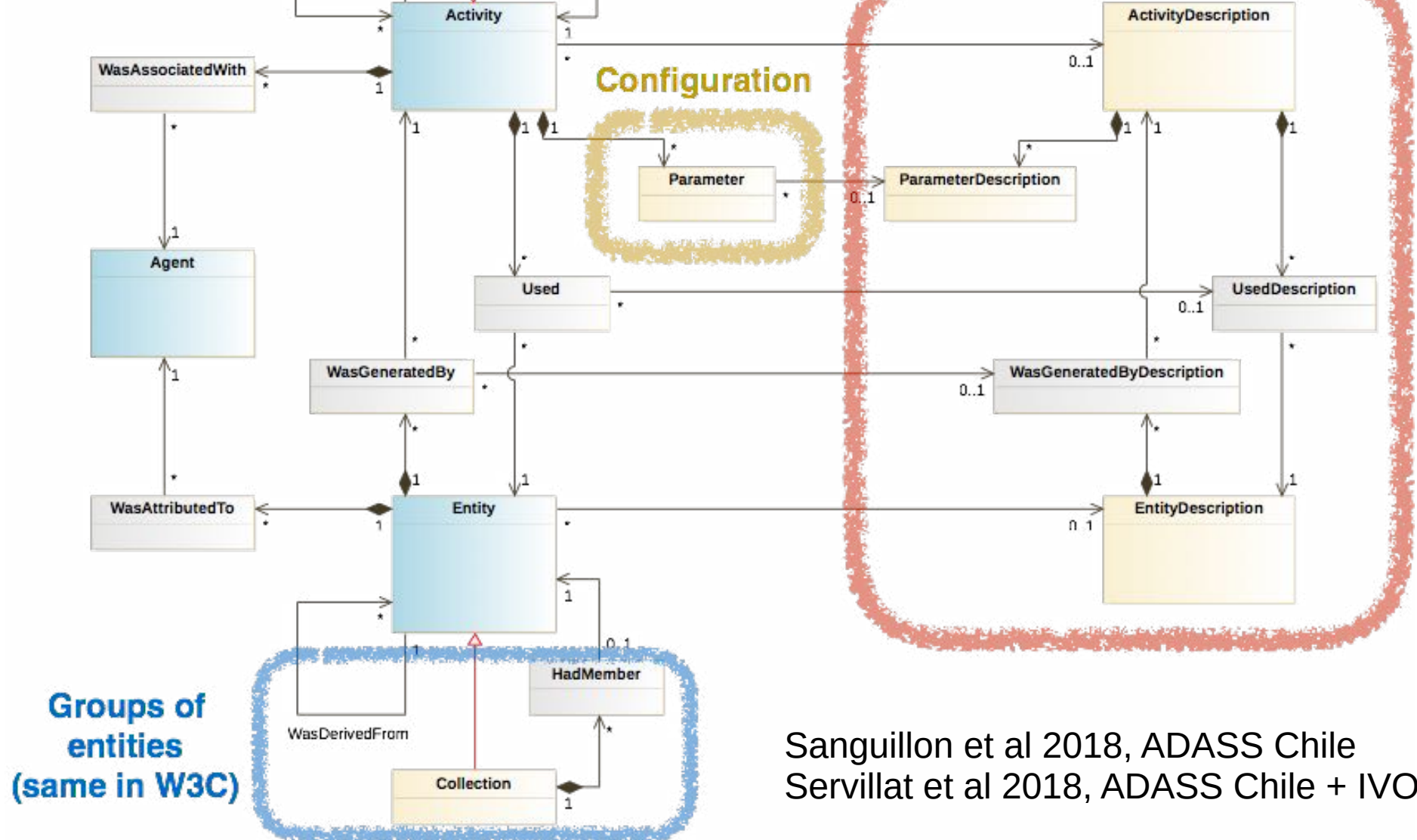
Editor(s)

Kristin Riebe, Mathieu Servillat

See presentation by M. Louys

Groups of activities

Description side



Sanguillon et al 2018, ADASS Chile  
Servillat et al 2018, ADASS Chile + IVOA









# Provenance during a CTA analysis step

- ◆ **OPUS** (Observatoire de Paris UWS Server) is a light **job controller** for the Paris Observatory **work cluster** developed in Python:  
<https://www.github.com/mservillat/OPUS>

OPUS Job Definition Job Manager Sign out admin

Job Description

Back to job list

Type	Start Time	Destruction Time	Phase	Details	Control
anactools_v1.1	2017-03-15 01:09:12	2017-04-14 01:09:08	COMPLETED	  	  

Job Properties

Job Parameters

Job Results

Job Details

- ◆ Follows the **IVOA UWS pattern**
- ◆ REST web service
- ◆ Job definition editor (ActivityDescription)
- ◆ Job manager
  - ◆ Stores job **properties** (start, stop time...)
  - ◆ **Parameters** also kept
  - ◆ Access to **results**
  - ◆ Visualization of **logs** and **Provenance information**

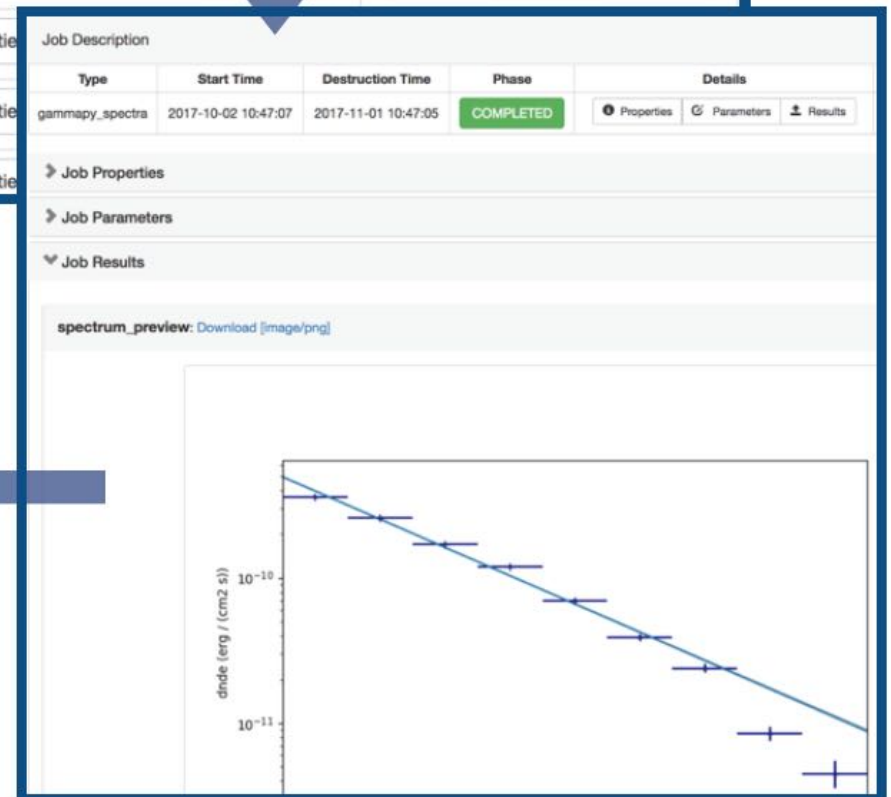
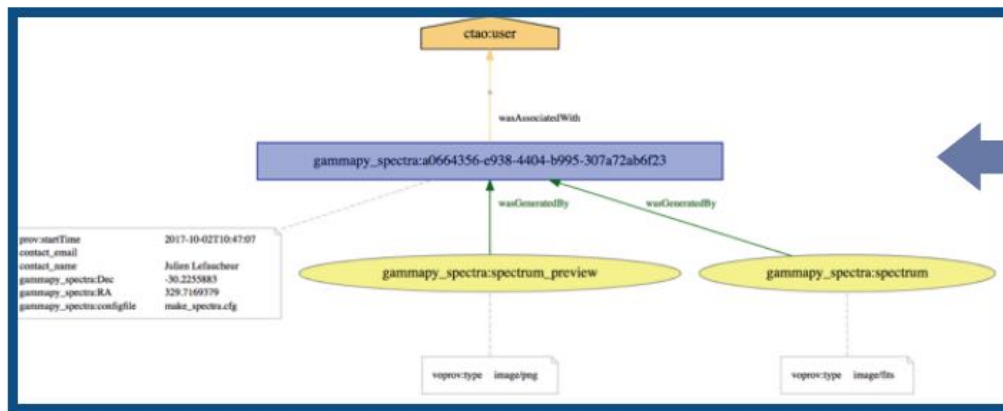
# From UWS to Provenance

OPUS Job Definition Job List Signed in as user

Job List for **gammapy\_spectra** Refresh Job List Create Test Job Create New Job

Type	Start Time	Destruction Time	Phase	Details	Control
gammapy_spectra	2017-10-02 10:47:07	2017-11-01 10:47:05	COMPLETED	Properties Parameters <b>Results</b>	Start Abort Delete
gammapy_spectra		2017-11-01 10:47:03	PENDING	Properties Parameters Results	Start Abort Delete
gammapy_spectra	2017-09-29 15:07:52	2017-10-29 15:07:51	COMPLETED	Properties	
gammapy_spectra	2017-09-29 14:55:10	2017-10-29 14:55:09	ABORTED	Properties	
gammapy_spectra	2017-09-29 14:21:20	2017-10-29 14:21:19	COMPLETED	Properties	

## Tracking of provenance information



<https://github.com/mservillat/OPUS>

# Gammapy – An open source Python package for gamma-ray astronomy

J. Lefaucheur<sup>1</sup>, C. Deil<sup>2</sup>, A. Donath<sup>2</sup>, L. Jouvin<sup>3</sup>, B. Khélifi<sup>3</sup> and J. King<sup>2</sup>  
<sup>1</sup>LUTH, <sup>2</sup>MPIK, <sup>3</sup>APC



MAX-PLANCK-INSTITUT  
FÜR KERNPHYSIK  
HEIDELBERG

l'Observatoire  
de Paris



## Context

- The current experiments (H.E.S.S., MAGIC and VERITAS) using the Imaging Atmospheric Cherenkov Telescopes (IACTs) technic can detect gamma-rays above a few dozen of GeV
- Data and tools for their analysis are private in the IACT community. The upcoming of the open observatory Cherenkov Telescope Array (CTA) slowly begins to change the mindset of the community
- Gammapy can be used to measure source properties such as morphology, spectrum and variability, using event lists as well as instrument response function (IRF) by taking into account IACT analysis methods' specificities

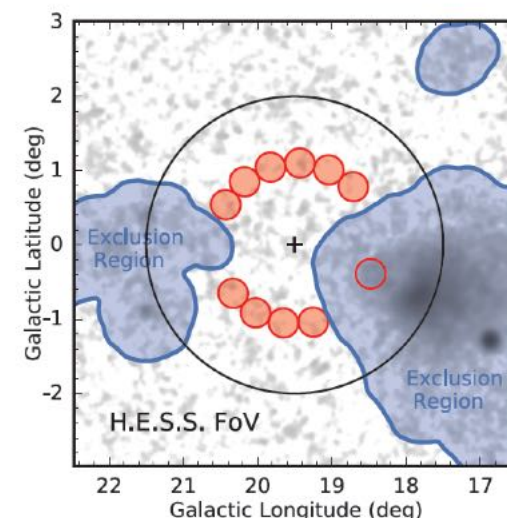
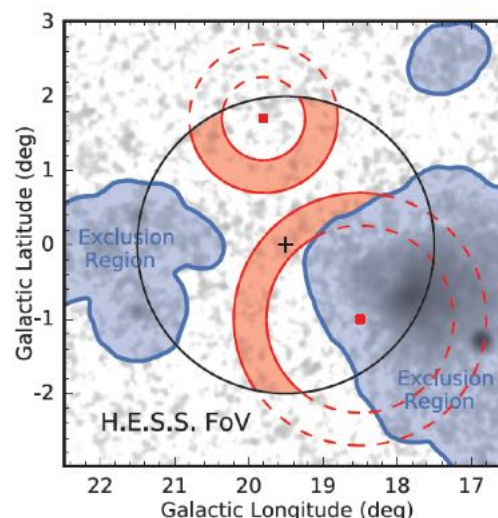


Figure 2: Illustration of classical method used to estimate the background implemented in Gammapy. The center of the field of view (FOV) is indicated with a black cross. Left, ring-background model used to reconstruct the morphology of a source. A ring (filled red regions) is used around a trial position (red squares) to estimate the background contamination. Right, reflected-region-background model used to reconstruct the spectrum of a source. The OFF regions (filled red circles) are used to estimate the background in the ON region (empty red circle). See Berge et al. 2007 for a detailed discussion.

Lefaucheur et al. 2018, ADASS Chile

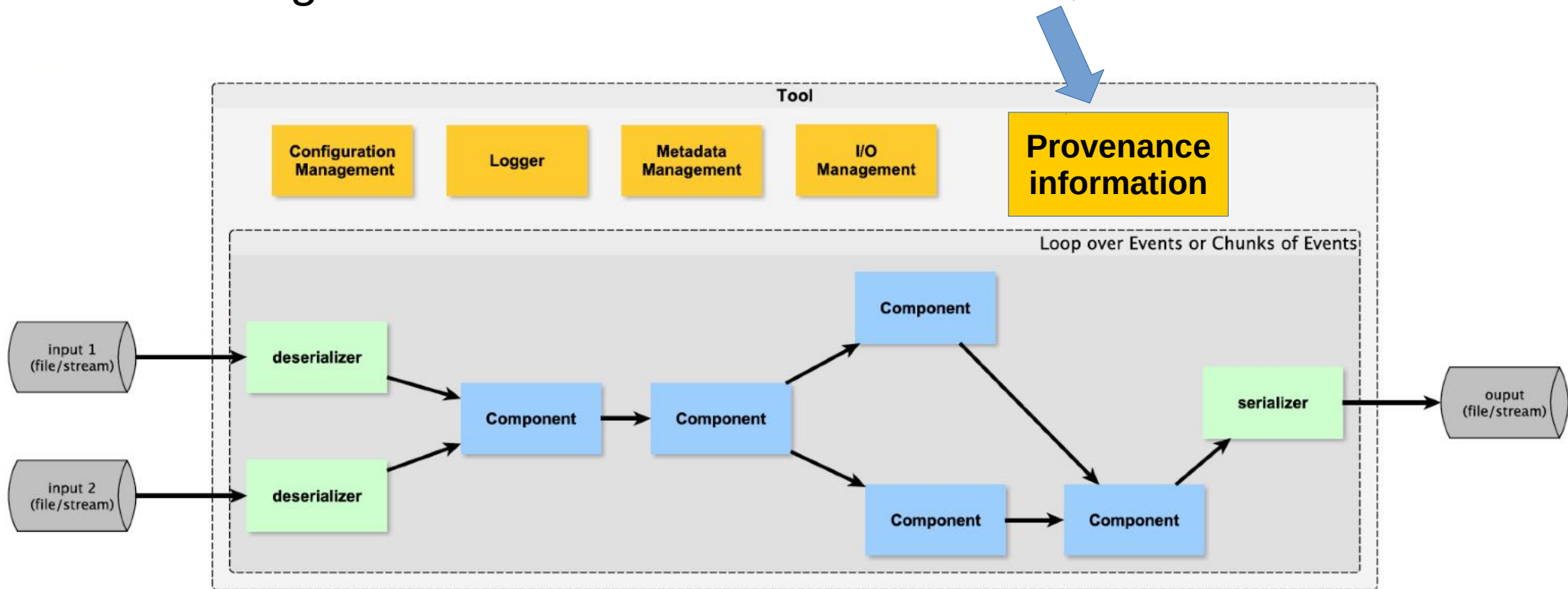


# Provenance in the Pipeline



cherenkov  
telescope  
array

- ◆ **Ctapipe**: a CTA data processing framework (prototype, not official, not recommended for use!)  
<https://github.com/cta-observatory/ctapipe>
- ◆ **Tool Python class** providing configuration, logger, I/O management... and **Provenance information**



@ Karl Kosack

# Provenance class for ctapipe

```
from ctapipe.core import Provenance

prov = Provenance()
# prov a singleton, so this gives you the same provenance class

prov.start_activity("some_activity")

... # do things
prov.add_input_file("test.txt")
prov.add_output_file("out.txt")

prov.start_activity("some_sub_activity")

# do more things
prov.add_output_file("out2.txt")

prov.finish_activity() # finish some_activity
prov.finish_activity() # finish some_sub_activity
```

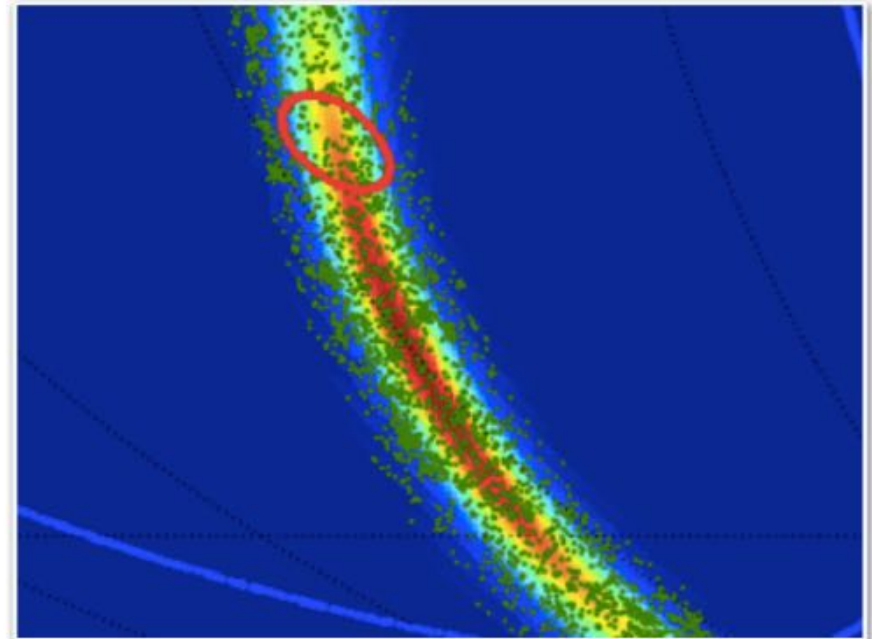
- ◆ Importance of **persistent identifiers**
- ◆ Also records **system configuration, state, and software versions**

# VOEvent for MW/MM CTA science cases

see Transient Alert Mechanisms workshop, Amsterdam, 2017  
<https://indico.astron.nl/internalPage.py?pageId=5&confId=62>

## Example: Processing of GW alerts

- IACTs are part of EM GW follow-up efforts
- Large uncertainty regions make follow-up challenging
- Advantages of H.E.S.S.
  - rapid slewing
  - relatively large FoV
- dedicated algorithms to determine optimized scheduling
  - 3D-correlation with galaxy catalog (GLADE) vs. 2D coverage of GW uncertainty region
- running fully automated within VO system
- decision on event-by-event basis

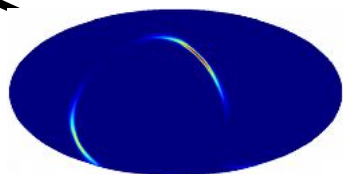


M. Seglar-Arroyo + FS, arXiv:1705.10138

@ Fabian Schussler



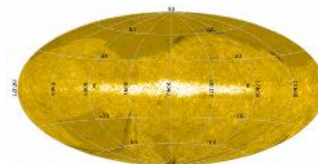
# Gravitational Waves: follow-up strategy



3D uncertainty region



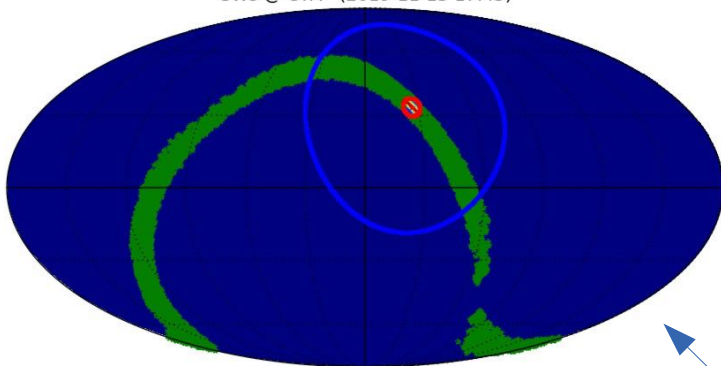
Galaxy catalog(s)



GW-Scheduler

pointings, priorities, etc.

GWs @ CTA (2019-11-15 17:45)



2.6184e-25 3.95212e-05

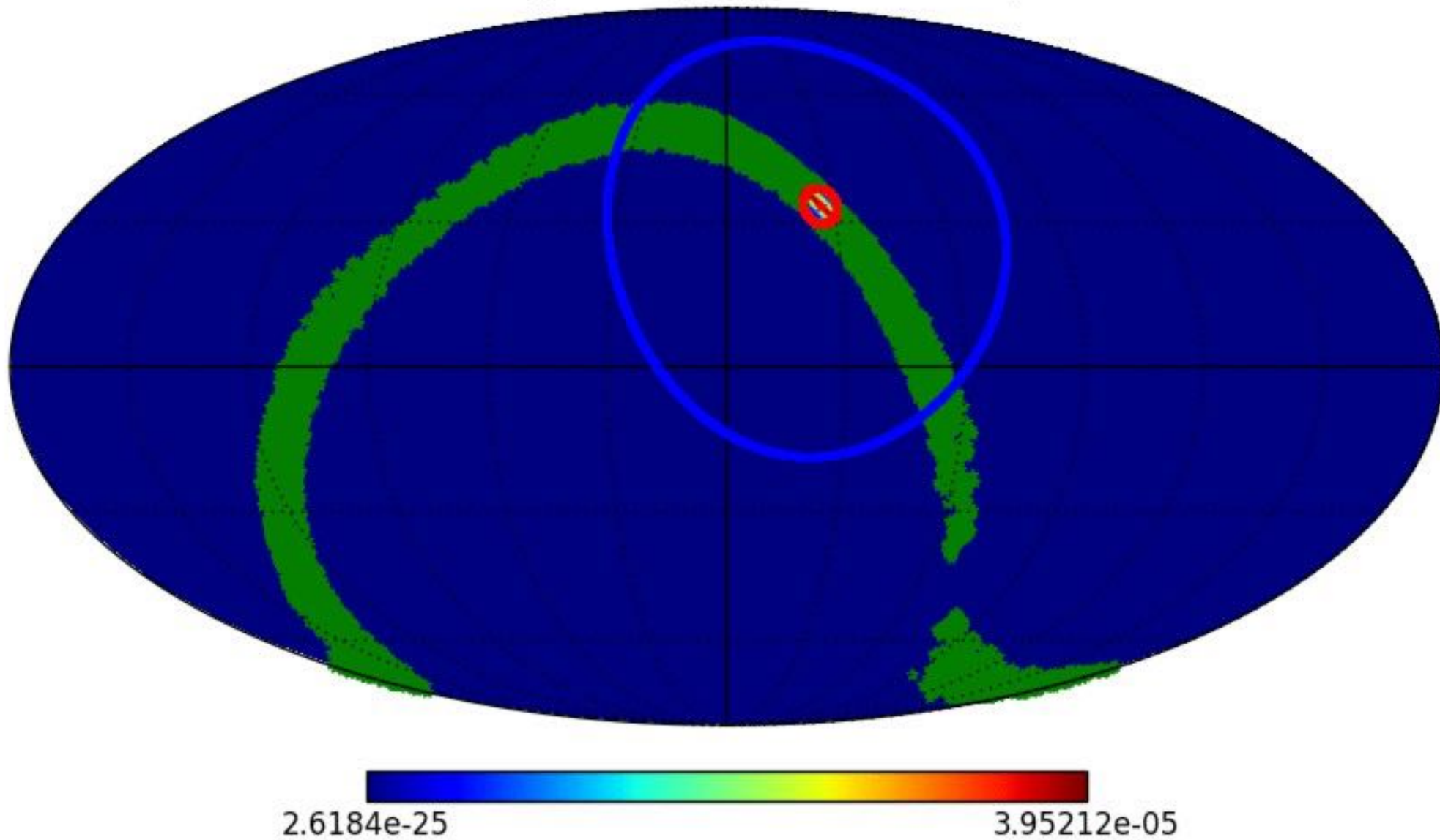
Blue rings : > 45° visibility N & S  
Green points : galaxies compatible with GW map  
Red : 8° CTA FoV

- **FULL CTA-ARRAY**
- Real-Time-Analysis searching for new/transient sources
- alert emission (internal/external)
- alert reception (EM counterpart)

simulated pointing strategy (worst case scenario: huge GW uncertainty)

@ Fabian Schussler

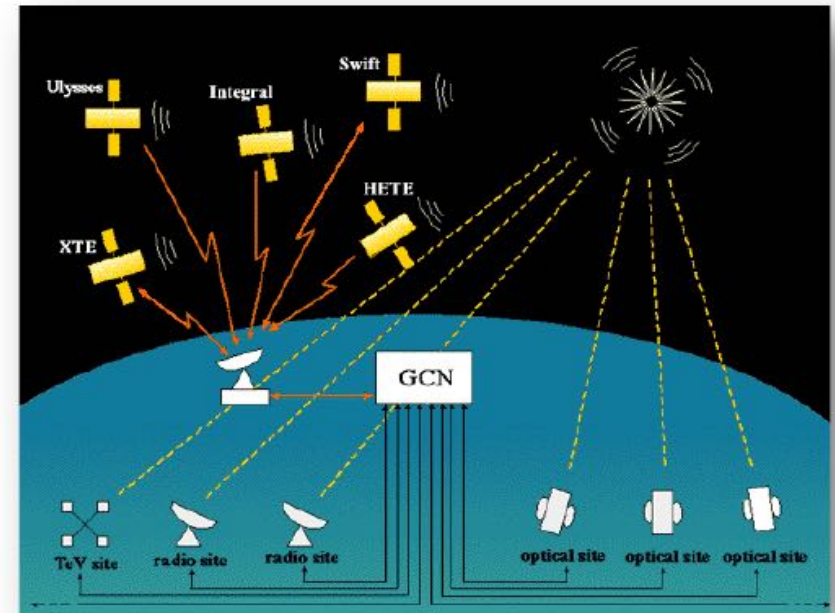
GWs @ CTA (2019-11-15 17:45)



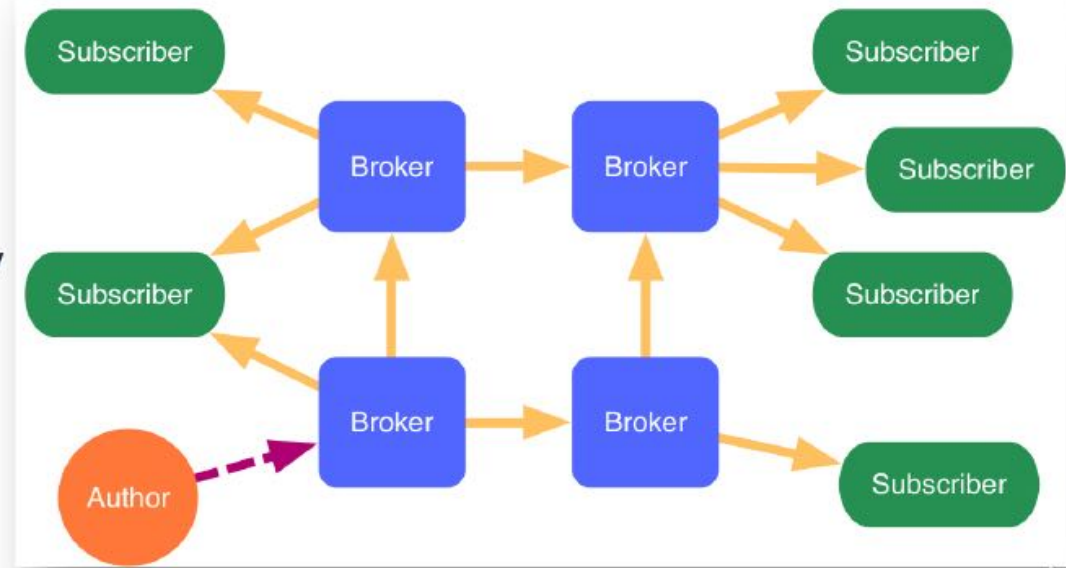
@ Fabian Schussler

# VO Event concept

- GCN Socket System
  - Hosted at NASA
  - Centralised
  - single point of failure
  - Can only select/deselect which packets to receive

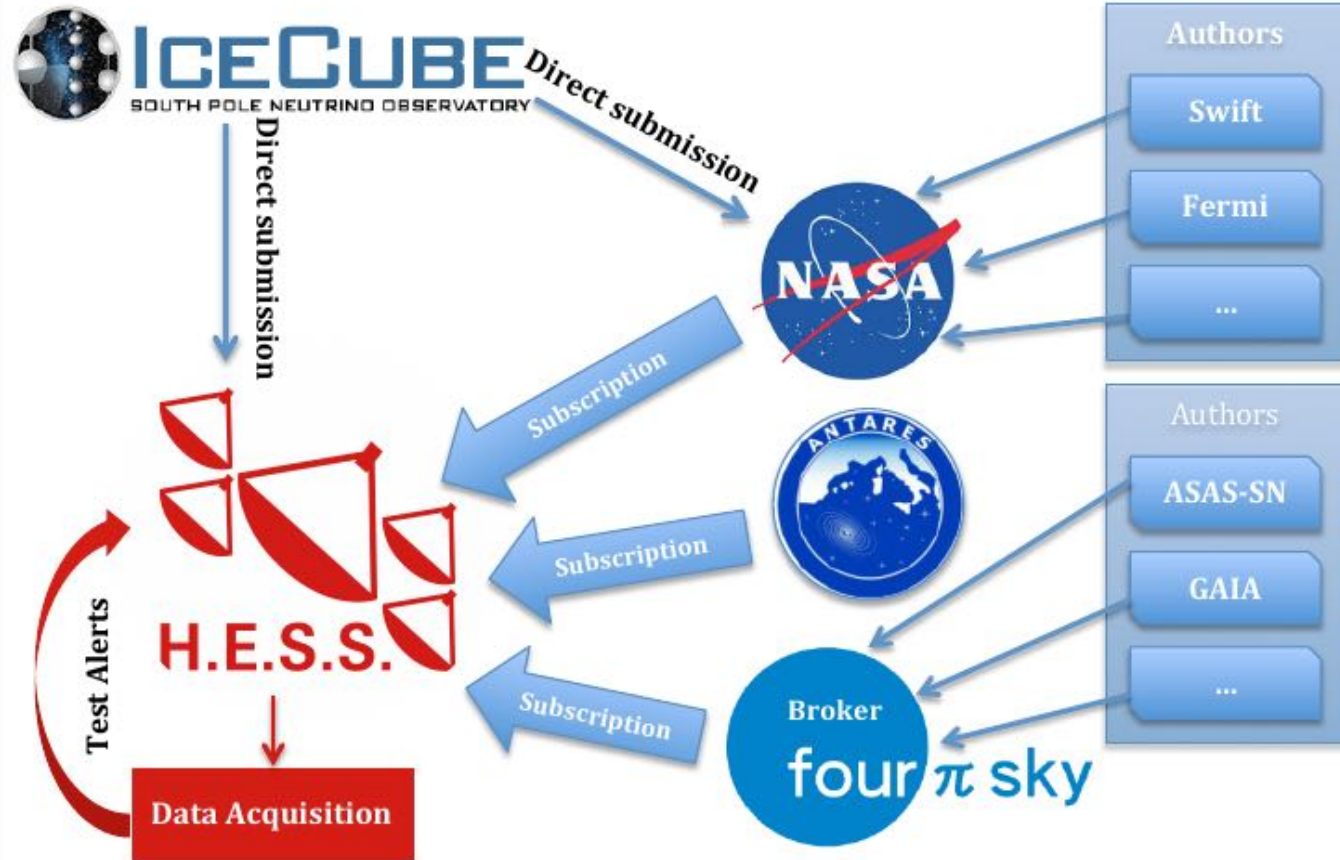
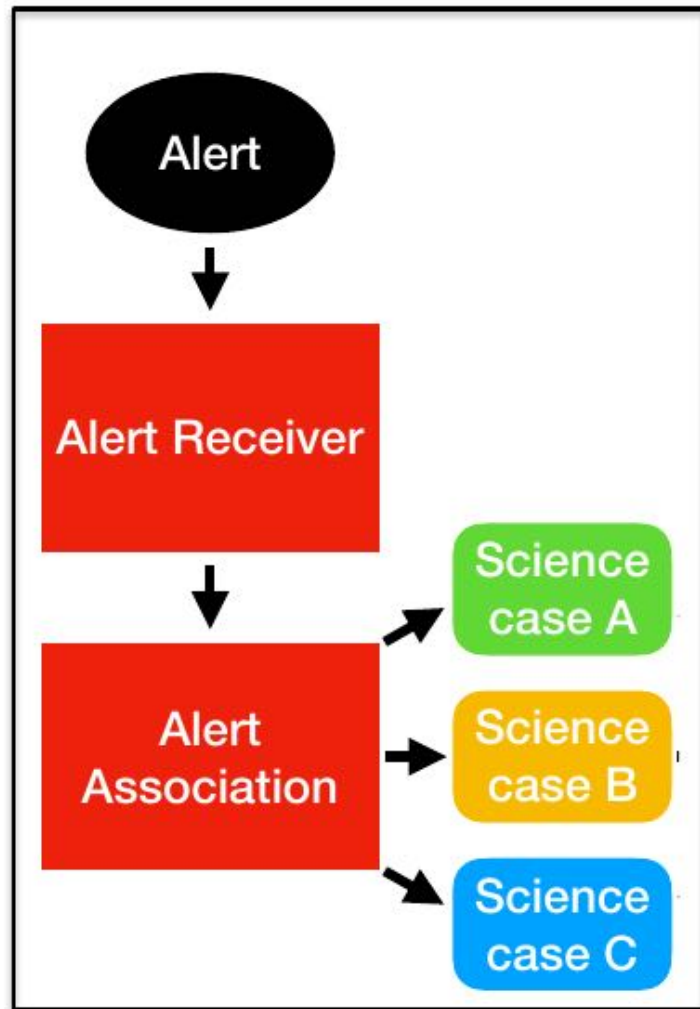


- VO Event System
  - Fully customizable setup
  - Distributed Broker system
  - Alerts can come from a variety of brokers
  - Feedback channels



@ Fabian Schussler

# The H.E.S.S. transient system – alert reception and association



@ Stefan Ohm



# ASTERICS Policy Forum – CTA document summary

	MW AGN & Transients	MM Campaigns	GP Transients
Physics Case	Flat spectrum radio quasars, BL Lacs, Radio galaxies, radio loud narrow line Seyfert 1 galaxies; prompt and afterglow emission in GRBs; FRBs	Gravitational waves and high-energy neutrinos from hadronic accelerators	Flaring pulsar wind nebulae, micro-quasars, pulsar binary systems, magnetars, novae, tidal disruption events, Galactic Centre, colliding wind binaries, and serendipitous source
Facilities	Monitoring facilities, SKA and precursors; AGILE, Fermi-LAT and X-ray satellites	SKA, ELT, GW second generation detectors: LIGO, VIRGO, KAGRA, I-LIGO, IceCube & KM3NeT, Baikal	Fermi-LAT, INTEGRAL IBIS, Swift BAT, eASTROGAM, XMM-Newton, Chandra, IXPE, XIPE, Athena, optical TLCs and SKA
Surveys/Time synchronisation	Monitoring and ToOs	Alerts of the order of seconds/minutes following CTA observations	Simultaneous X-ray (similar timescale variability); ToO observations for MW follow-ups and shared time
VOTools / Archival Data	Fundamental to obtain information about light curves and SEDs	Crucial for localisation and identification	Fundamental for source identification and observation strategy optimisation
Modeling/Numerical Simulations	3D MHD & GRB modeling	VHE predictions; high-performance computing	SED and MHD simulations
Requested Time	~100 h/yr	~50 h/yr	~150 h/yr